



FOCUS
Focal Plane Array for Universe Sensing



LABORATORIO NAZIONALE
ADONI
OTTICA ADATTIVA



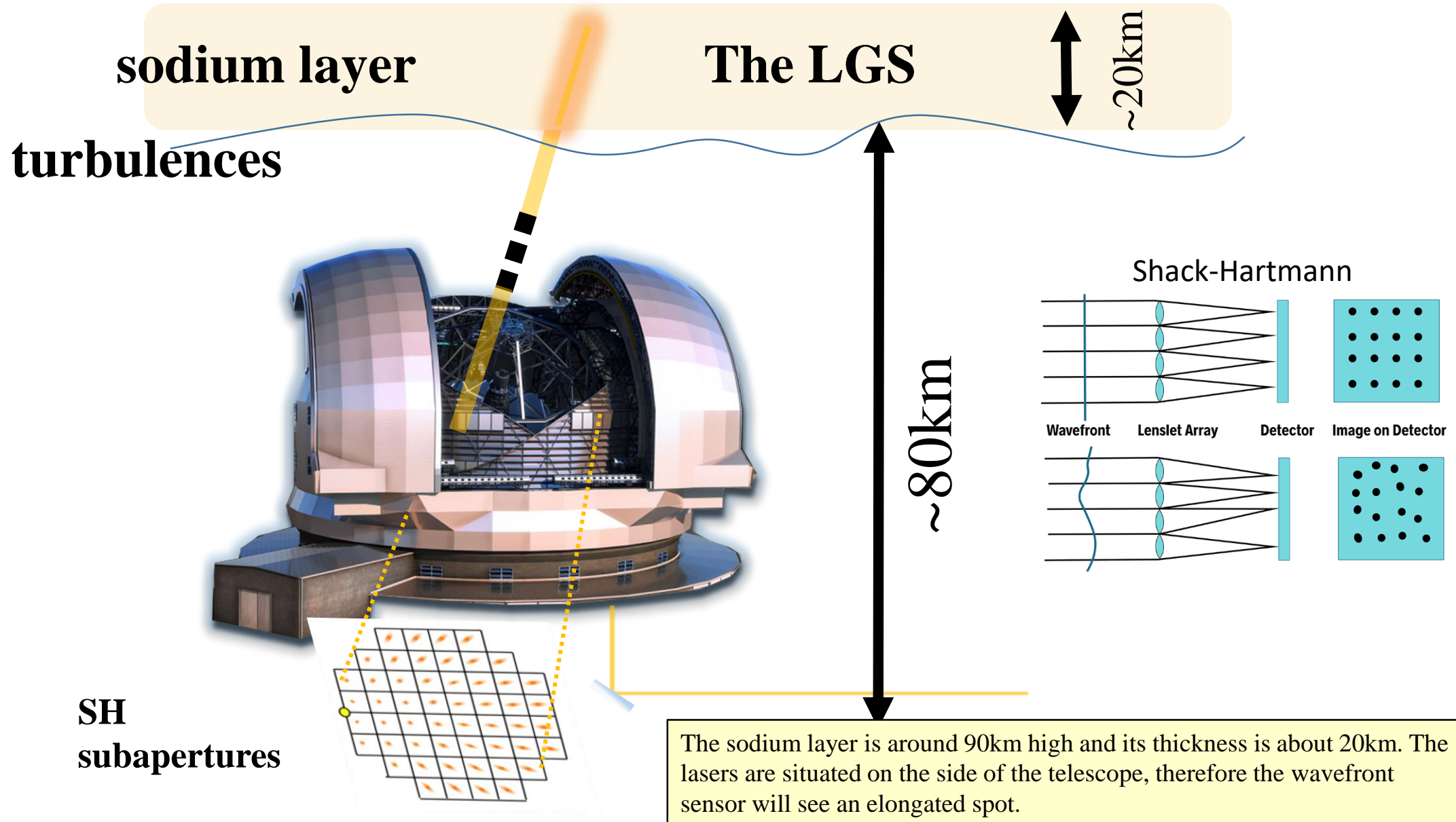
Aix-Marseille
université



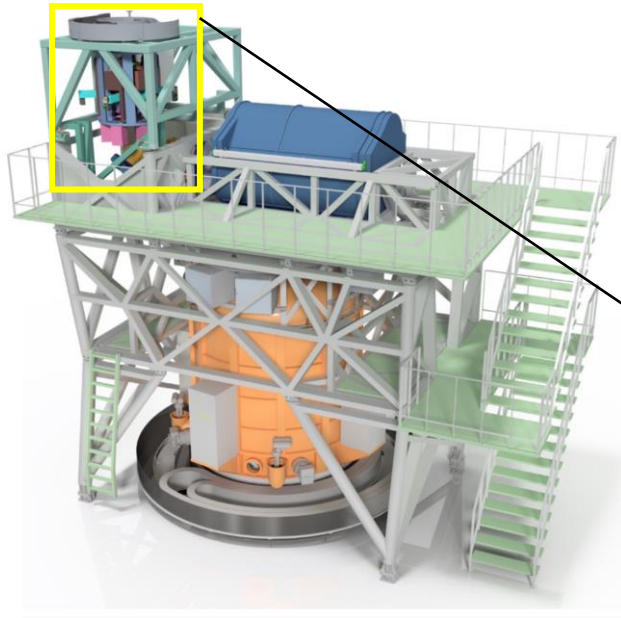
HARMONI at ELT: Full scale prototype of the LGS WFS

Pierre Jouve, Felipe Pedreros, Anne Costille, Kjetil Dohlen, Kacem El Hadi, Thierry Fusco, Zibo Ke, Benoit Neichel

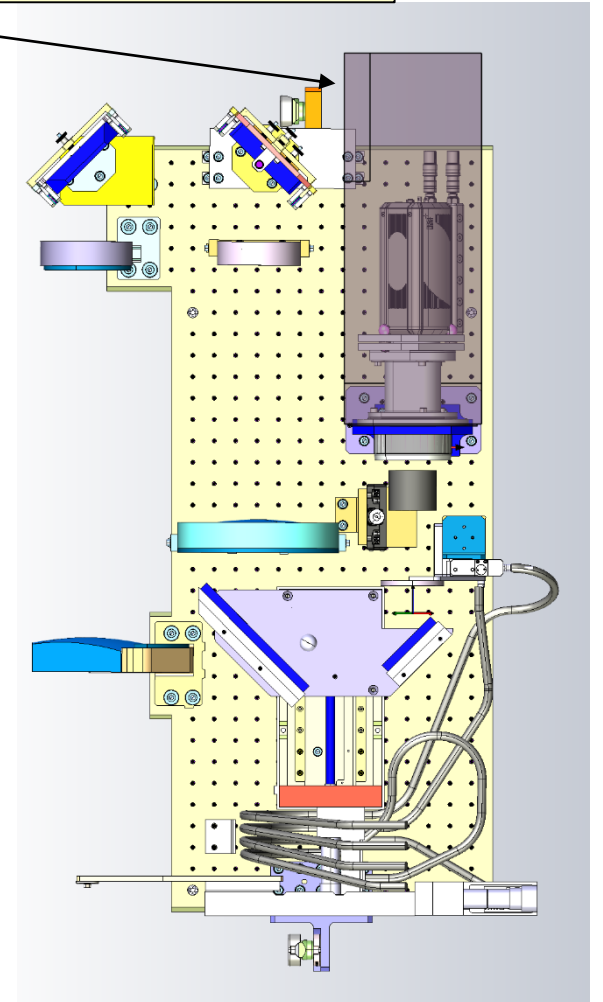
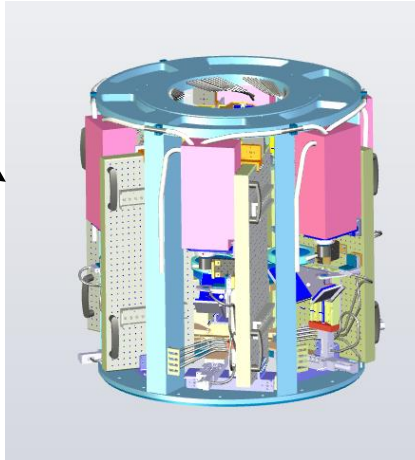




LGSS (Laser Guide Star Sensor)



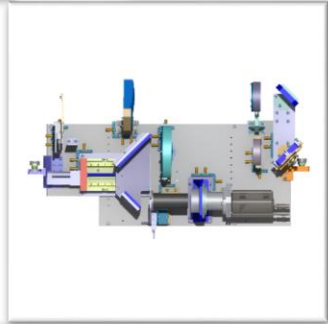
LDM (Laser guide star Detector Module)



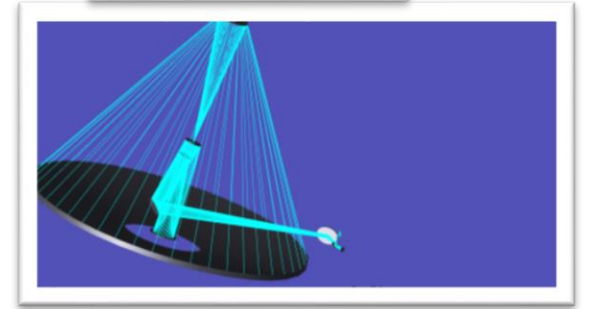
Goals for Harmoni:

- Physics of distant galaxies
- Stars in other galaxies
- Characterization of giant exoplanets
- Relationship between black holes and their galaxy
- ...

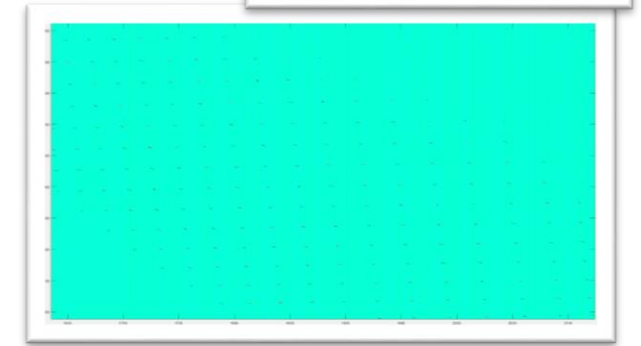
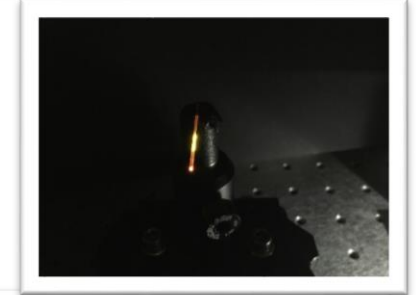
-Characterization of the WFS, camera, relay, microlenses array, SLM.

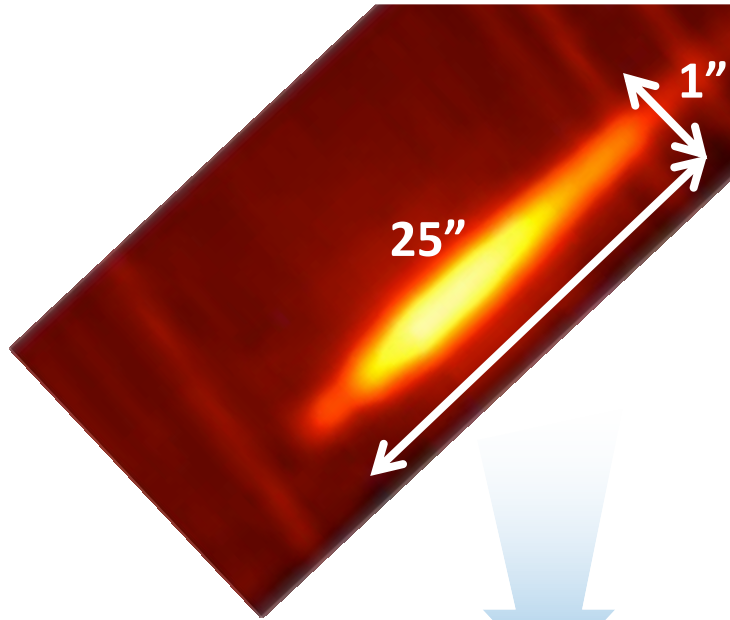


-Optical design to emulate the ELT and turbulences. Innovative solution for elongated source



-AO performances, closed loops, COGs vs WCoGs, interaction matrix analysis, IM simulations.





Ideally, we need subapertures with
25x25 pixels of ~1"/pixel

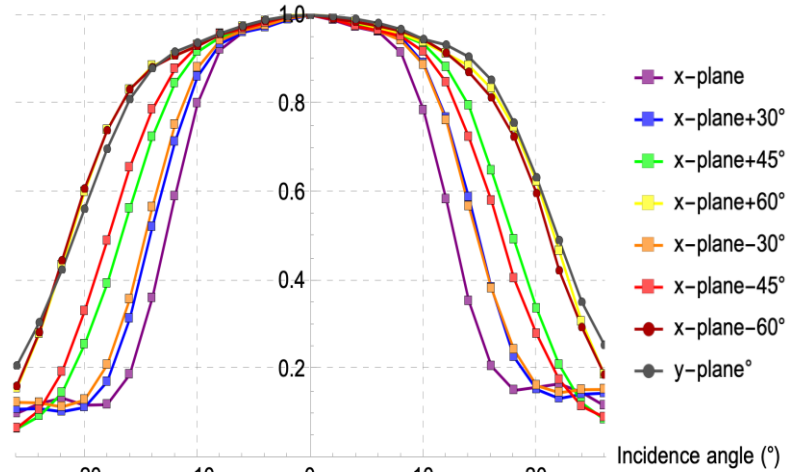
For 80x80 subapertures, we need
2000 x 2000 pixels



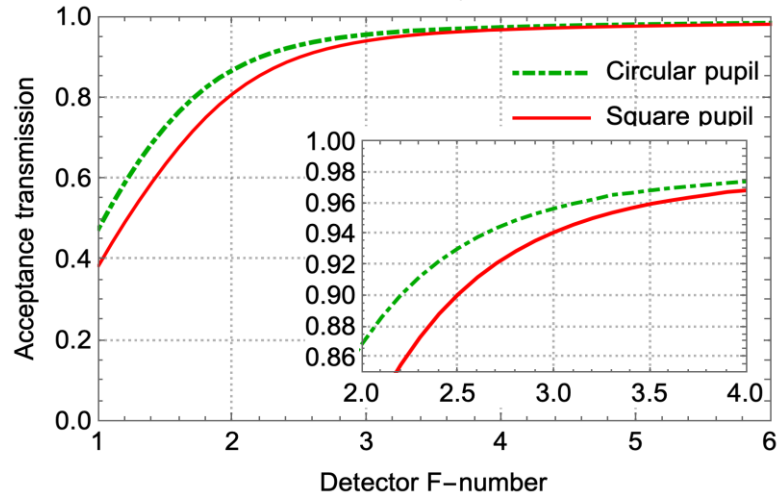
2000x2000 pixels detector, running at 500Hz,
with RON<3e- does not exist...

Instead, we are using a C-Blue camera with:
1608x1104 pixels
RON<3e-
500Hz
Global shutter

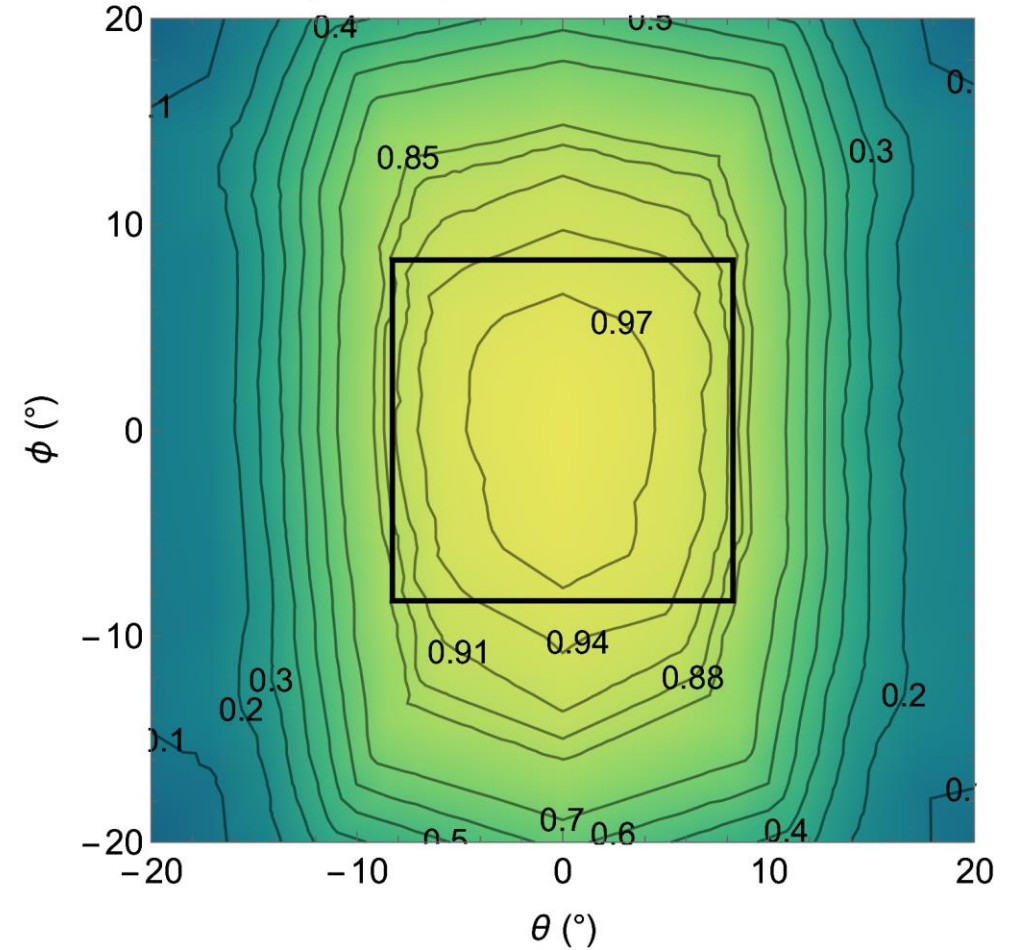
Measured angular pixel response (normalized), C-BLUE camera, $\pm 26^\circ$ incidence



Transmission efficiency, CBLUE camera

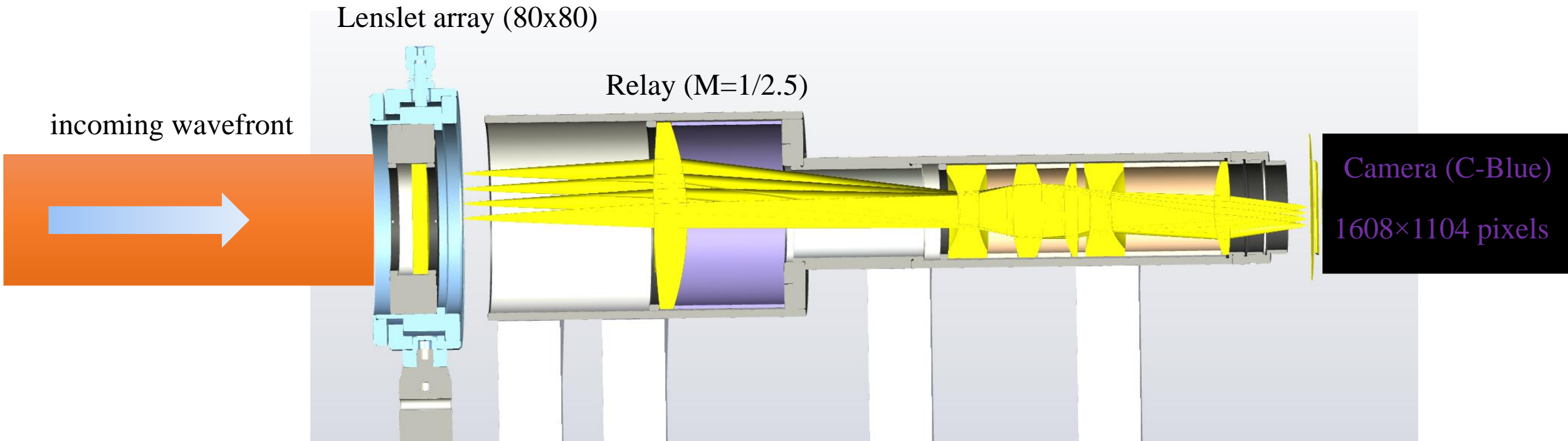


Pixel angular response function C-BLUE camera

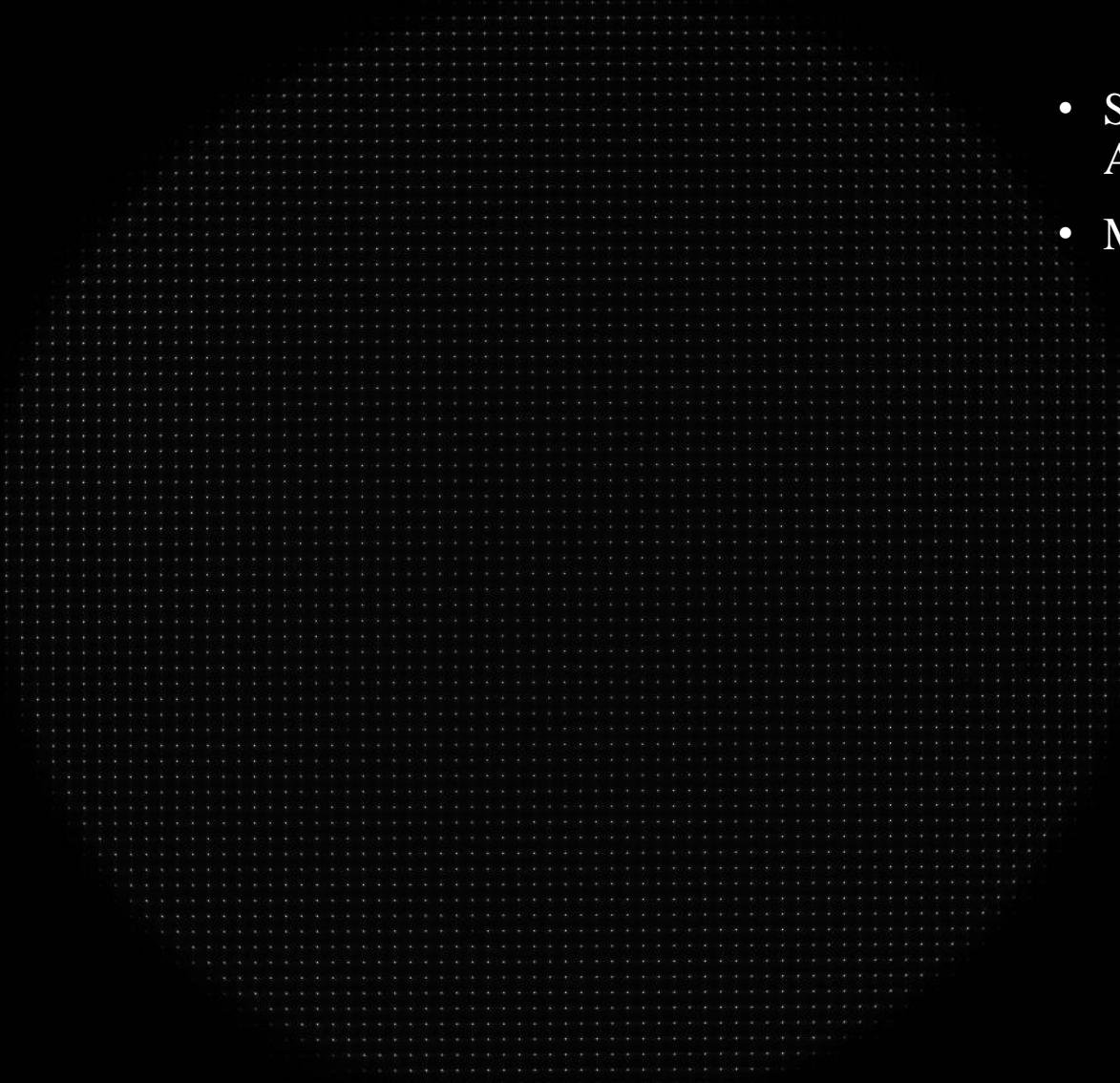


The LDM (Laser guide star detector module)

- The wavefront sensing is done with a **Shack-Hartmann**.
- The focal of the microlenses is small (large field 15", 80x80 subapertures to sample the 39m M1), therefore we use a relay to re-image the pupil on the camera.
- Due to the number of subpertures and the size of the elongated spots, we need an important number of pixels, therefore we are using a CMOS sony sensor 1608x1104 pixels. (C-Blue).

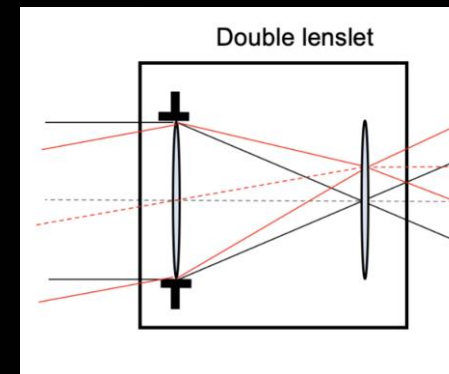


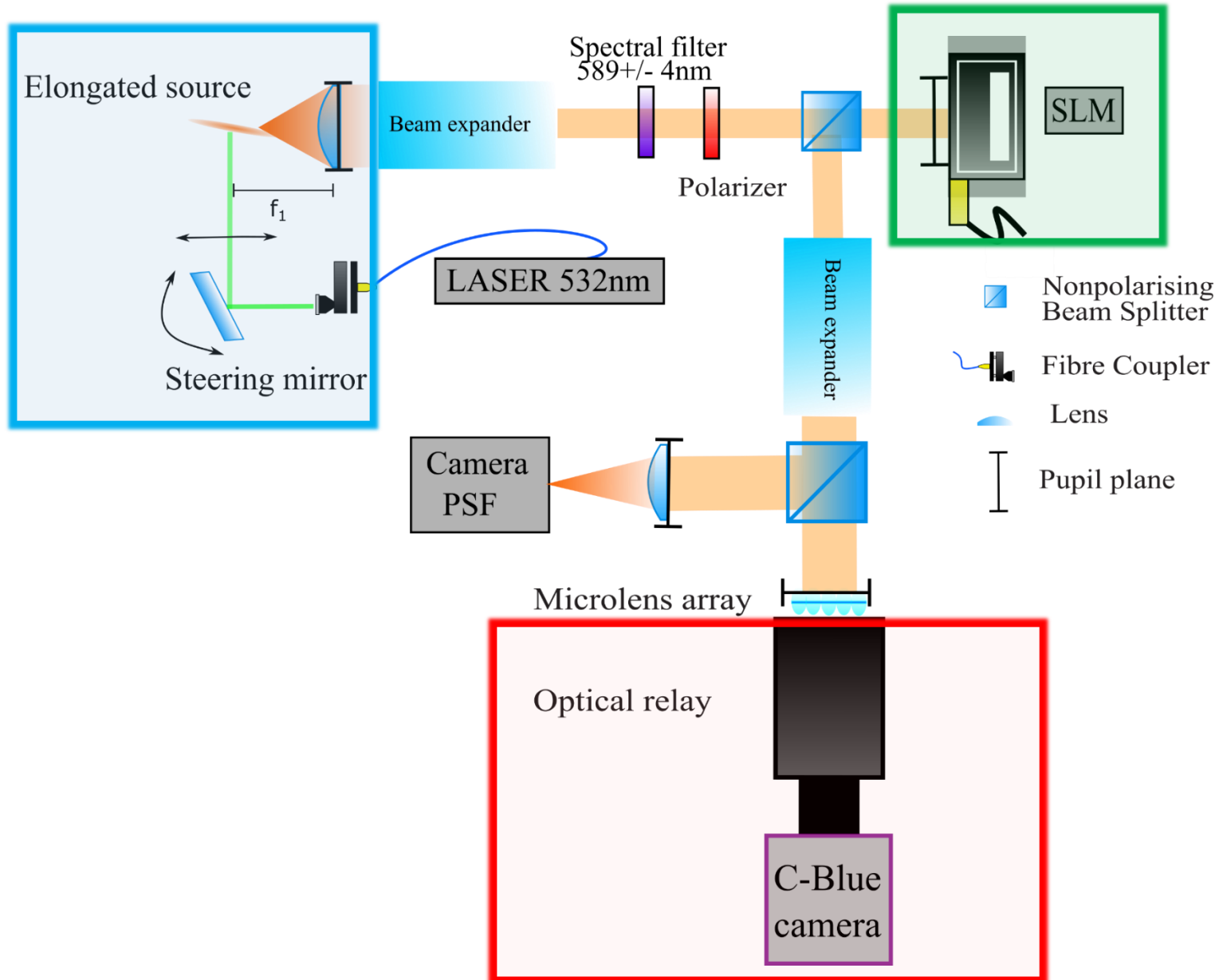
80 microlenses



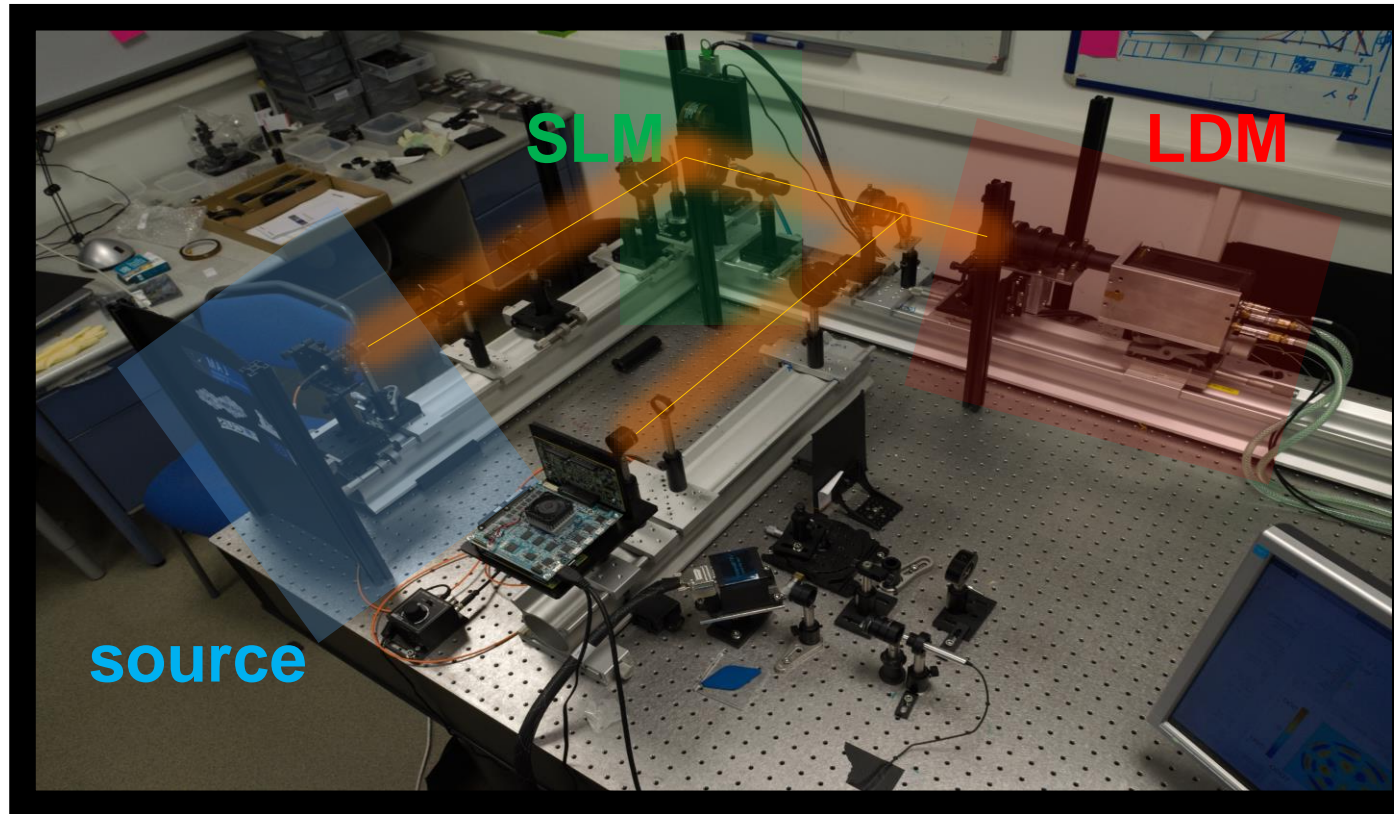
80 microlenses

- Specification of the microlens array manufactured by AMUS :
- Main characteristics:
 - Lenslet pitch : $290\mu\text{m}$
 - Lenslet focal distance : 2.5mm
 - Lenslet number: 80 (on circular aperture of 24.00mm diameter)
 - Number of « useful » lenslet : 80
 - Optimal pupil size on the lenslet array : 24.0mm
 - Lenslet F number : 8.645



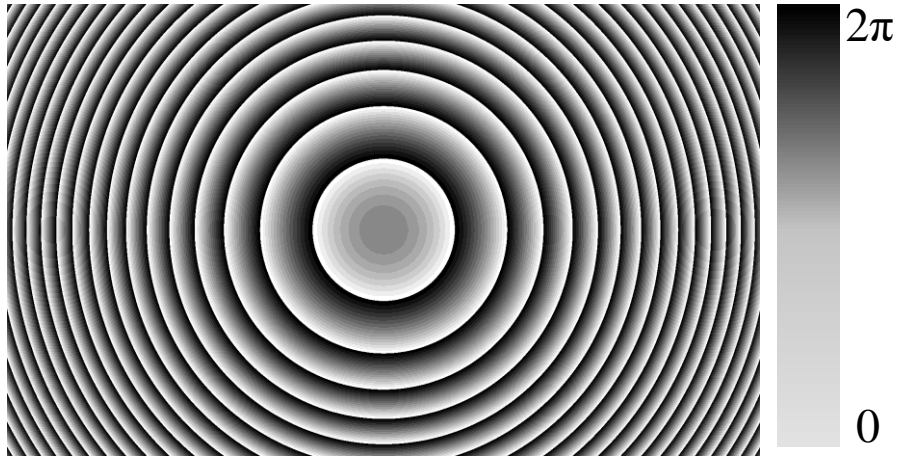


2nd step: we add phase aberation with a SLM



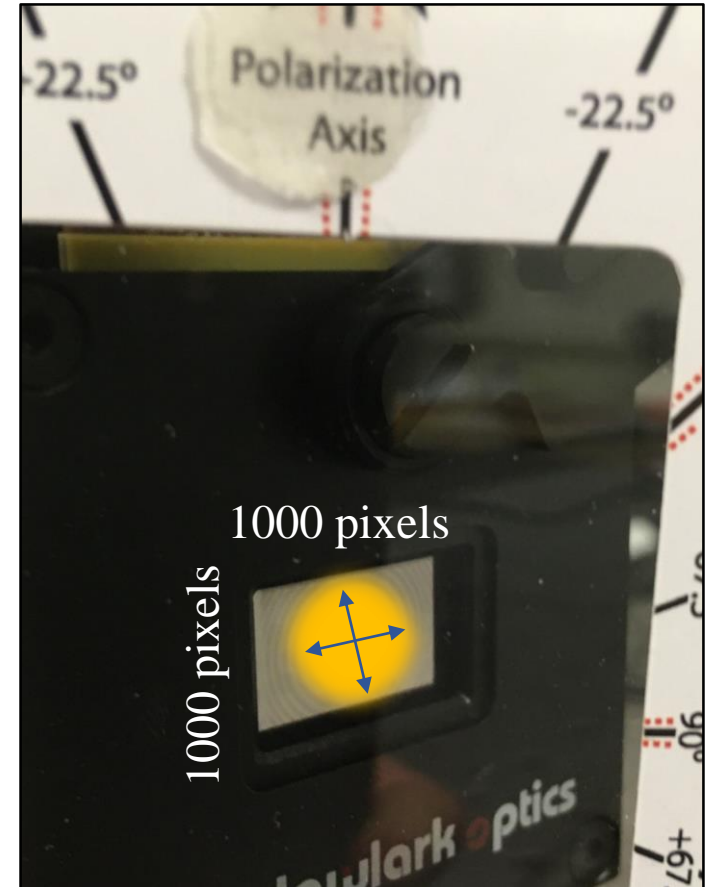
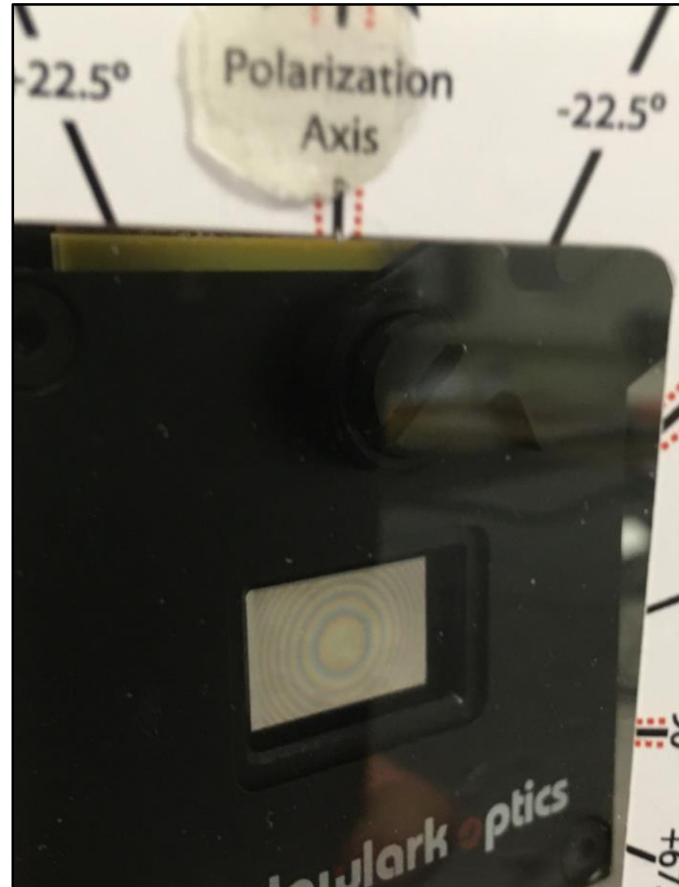
The SLM: Loading a Zernike phase map

Phase map loaded into SLM Focus

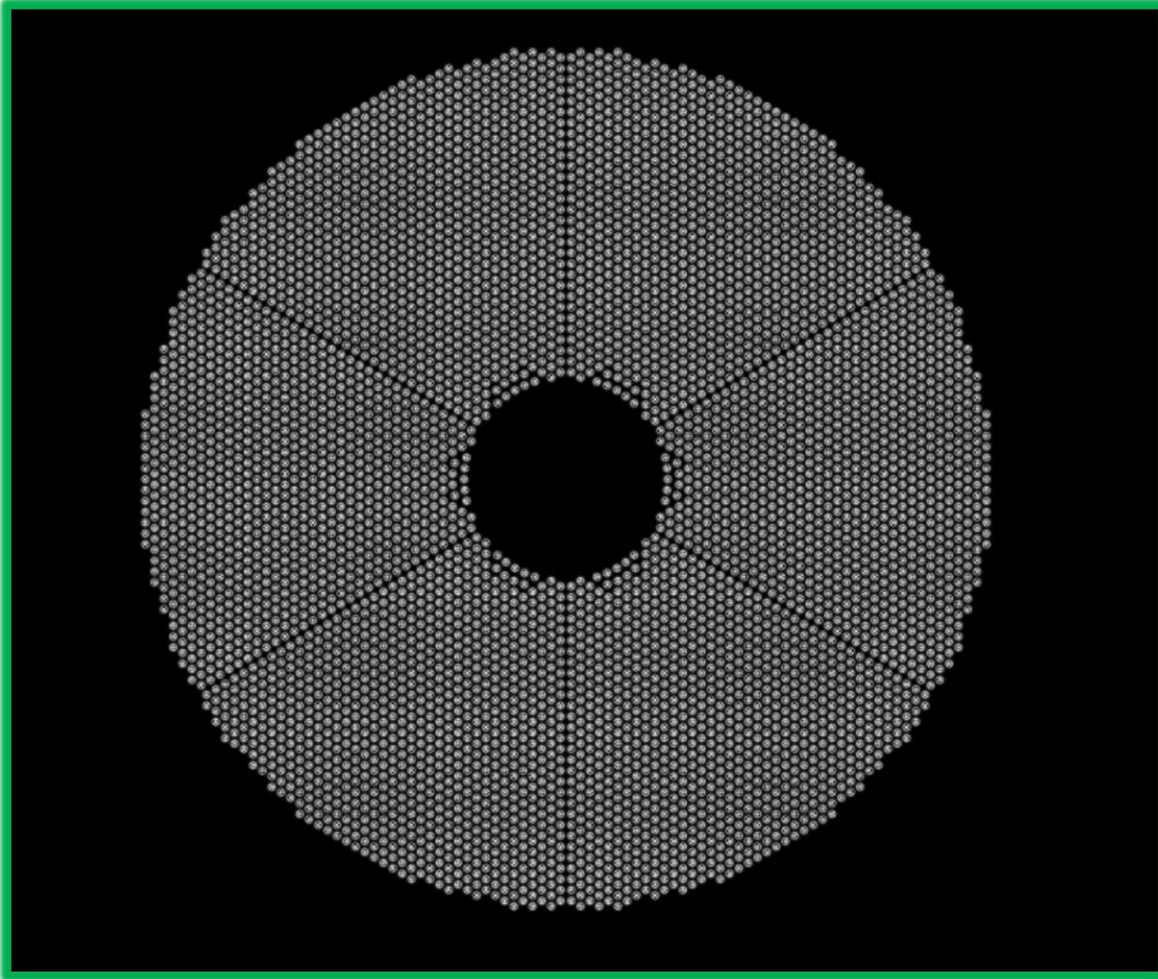


Polarization axis of liquid crystals
(aka *slow axis*, *extraordinary axis*) in
this SLM is **vertical**

SLM seen through polarizer film

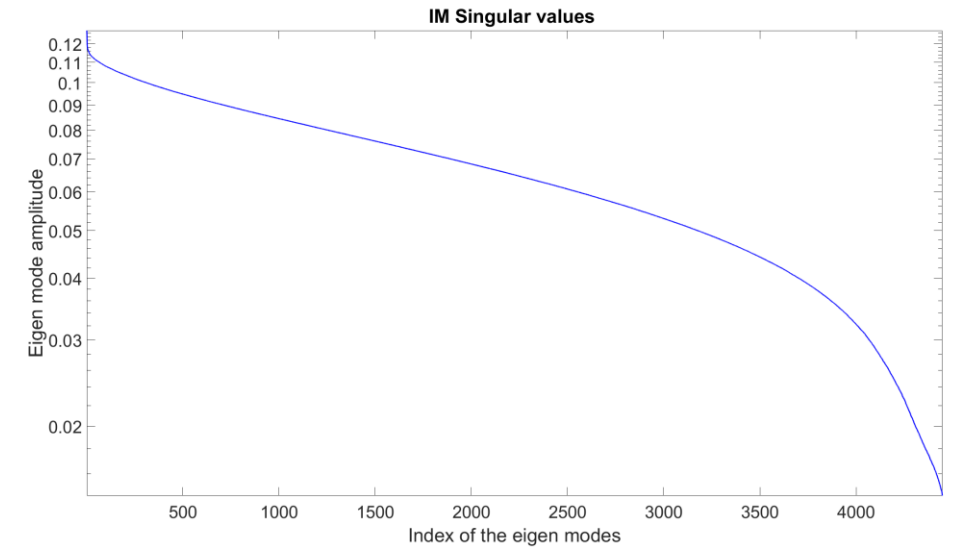
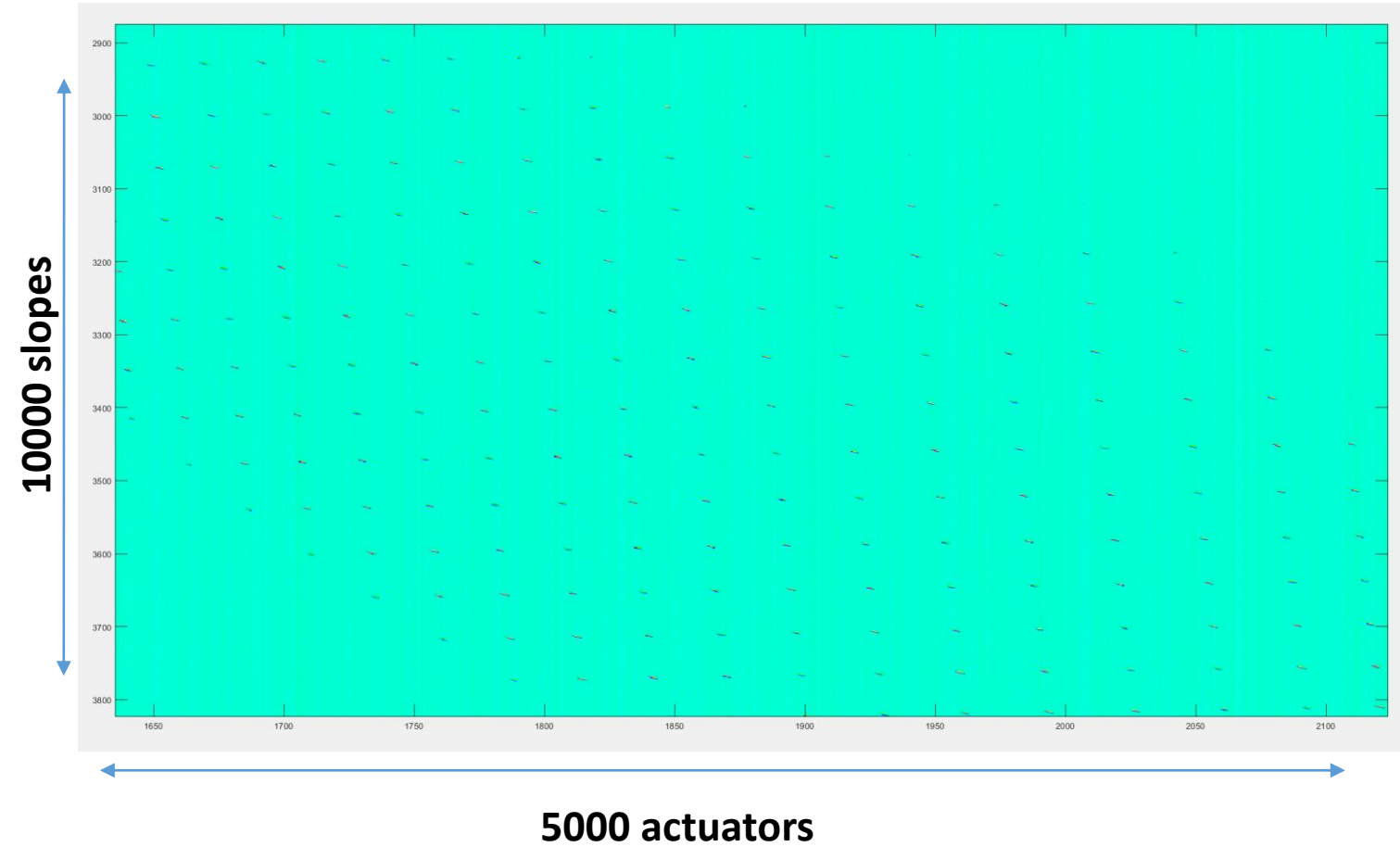


Vertical polarization



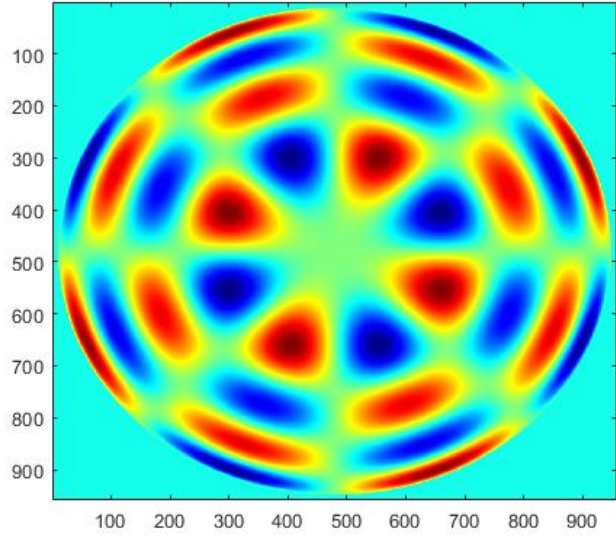
M4 influence fonctions projected on M1. We use that geometry to create interaction matrices. We also acquired IM in the Fried configuration, with gaussian as influence functions, with ESO influence functions. The SLM allows us to change the geometry or the influence function easily.

Experimental interaction matrix

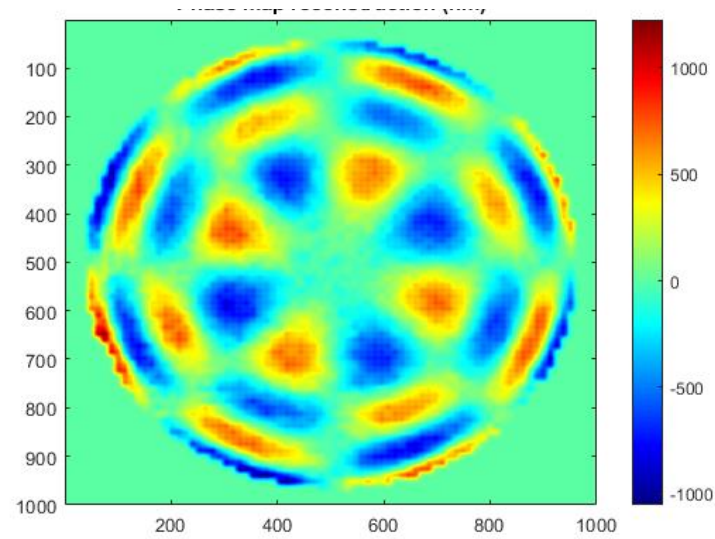


Experimental interaction matrix

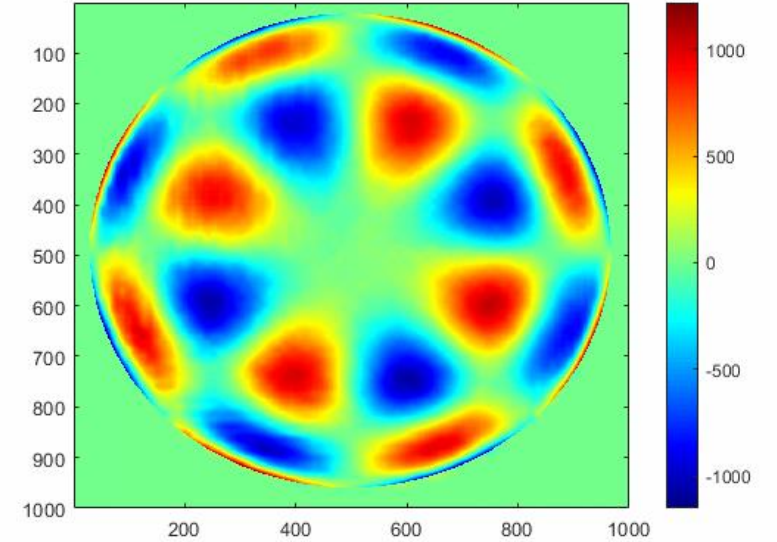
Phasemap printed on the SLM



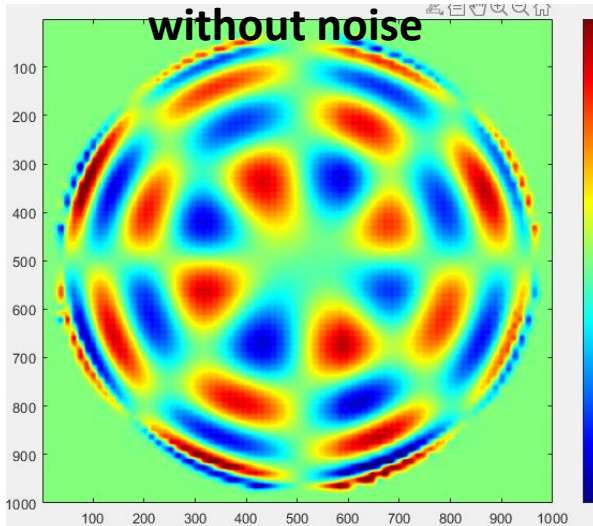
Experimental reconstruction



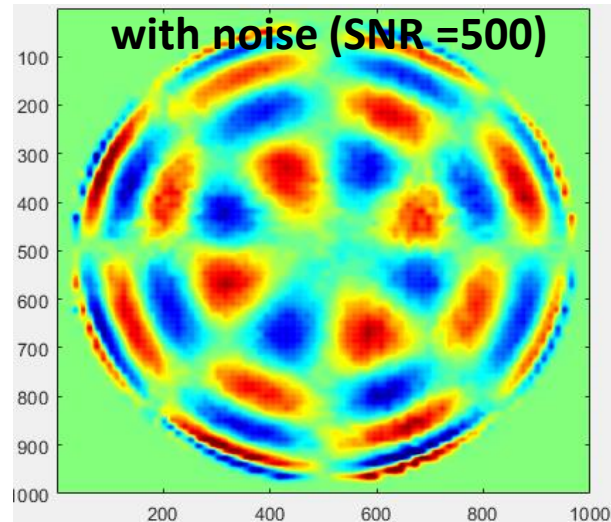
Closing loops experimentally



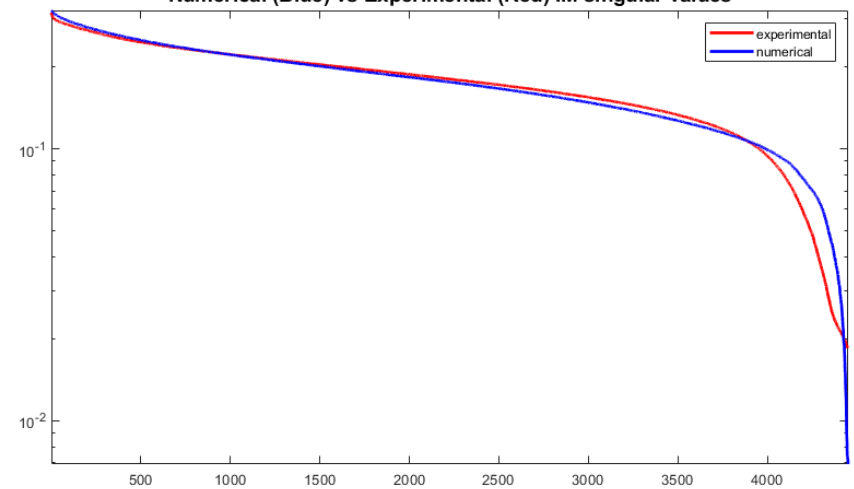
reconstruction simulation



reconstruction simulation



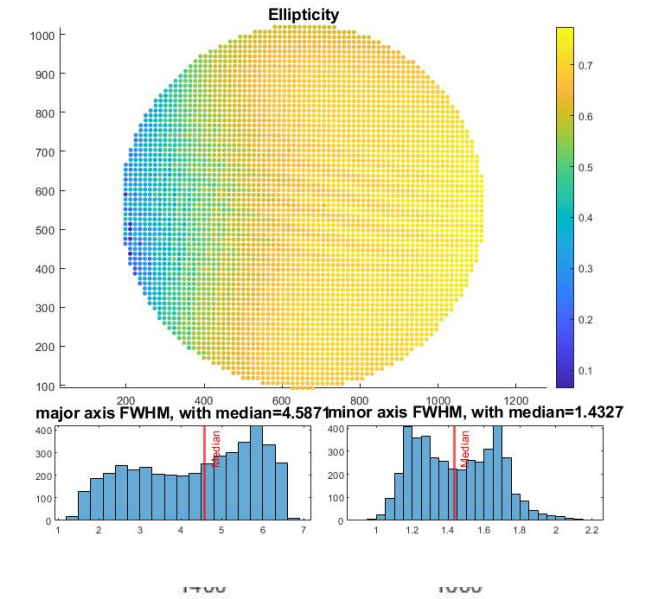
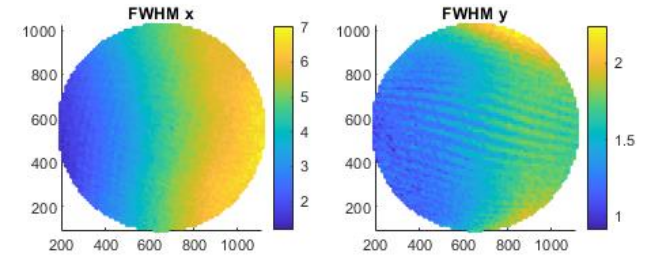
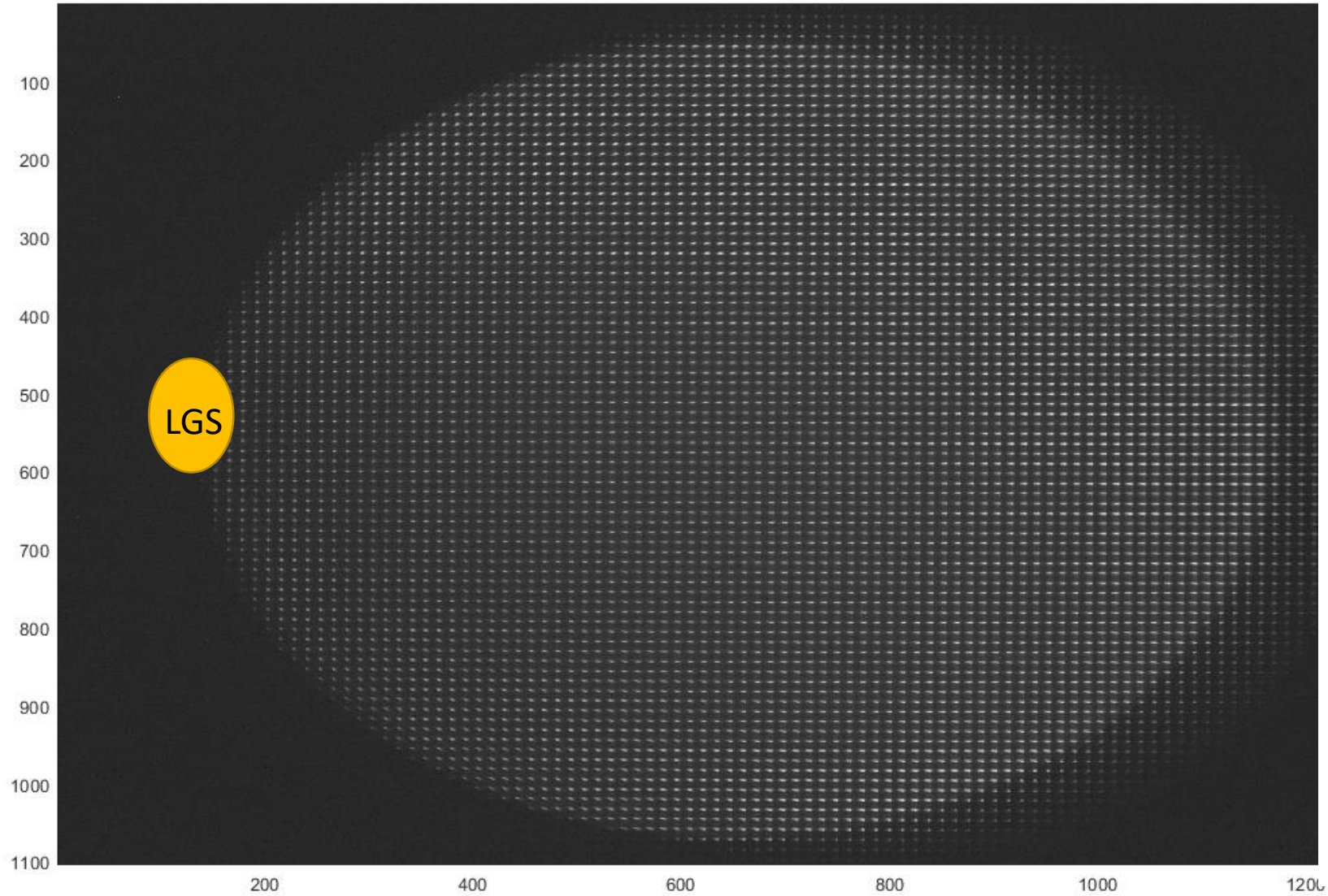
Numerical (Blue) vs Experimental (Red) IM singular values

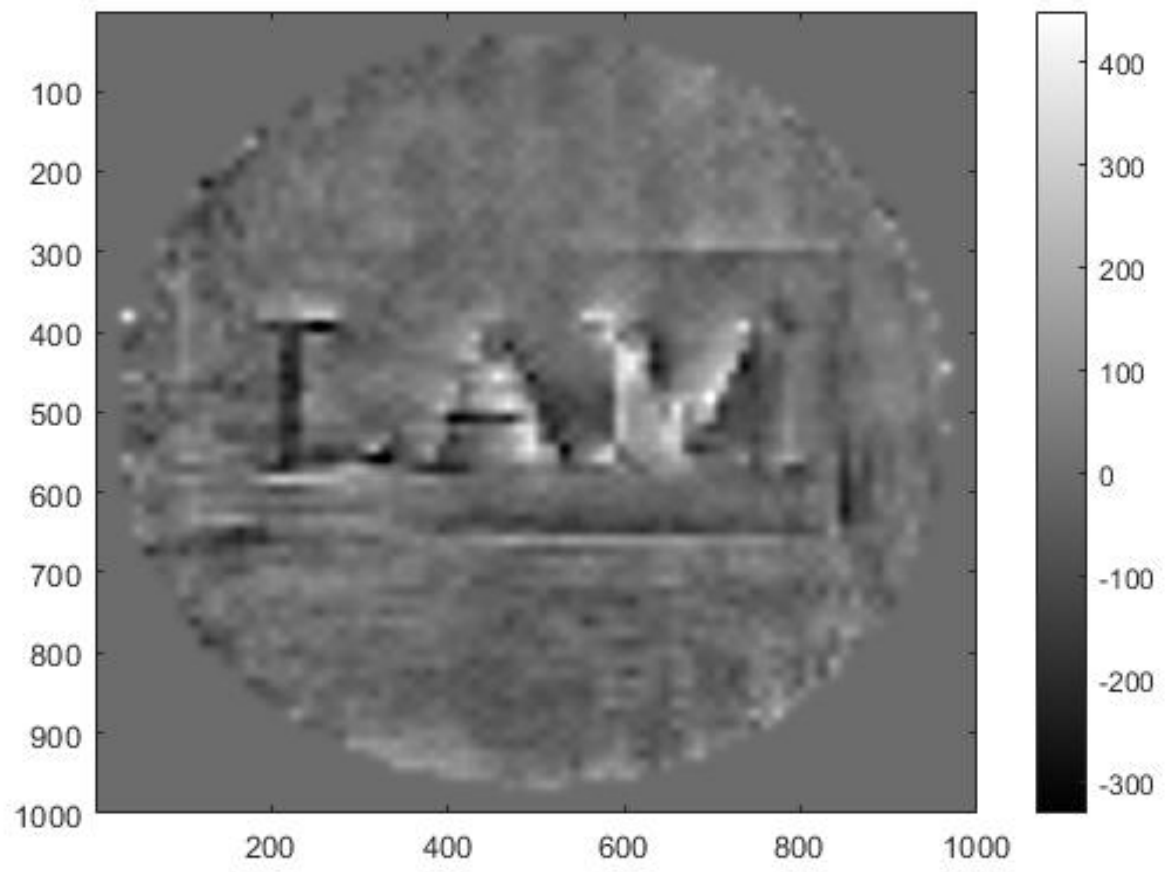


Next step: with elongated spot

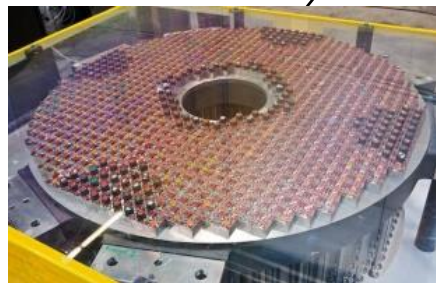
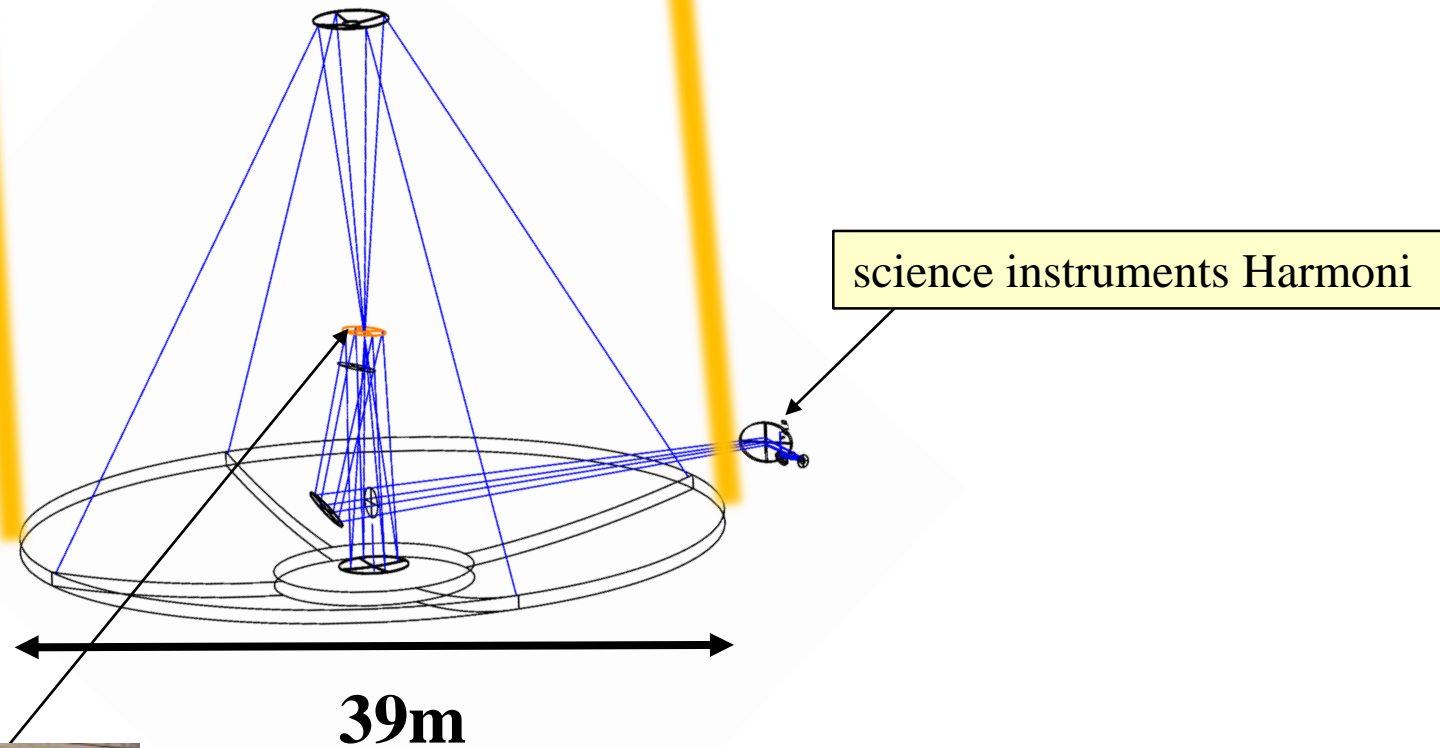


Next step: with elongated spot

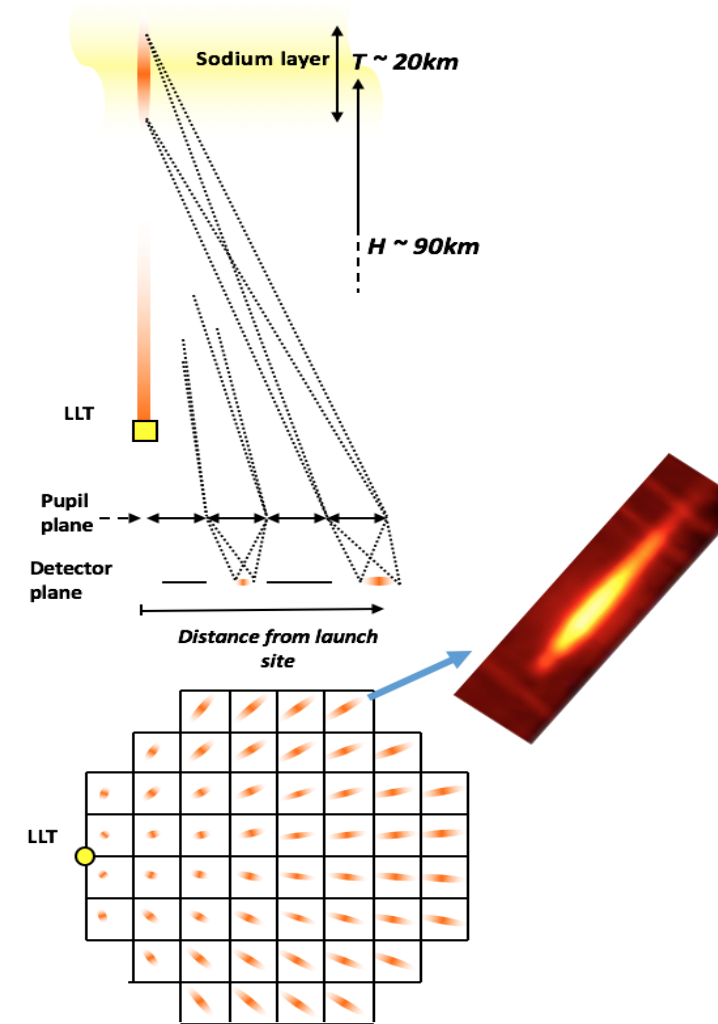




Merci

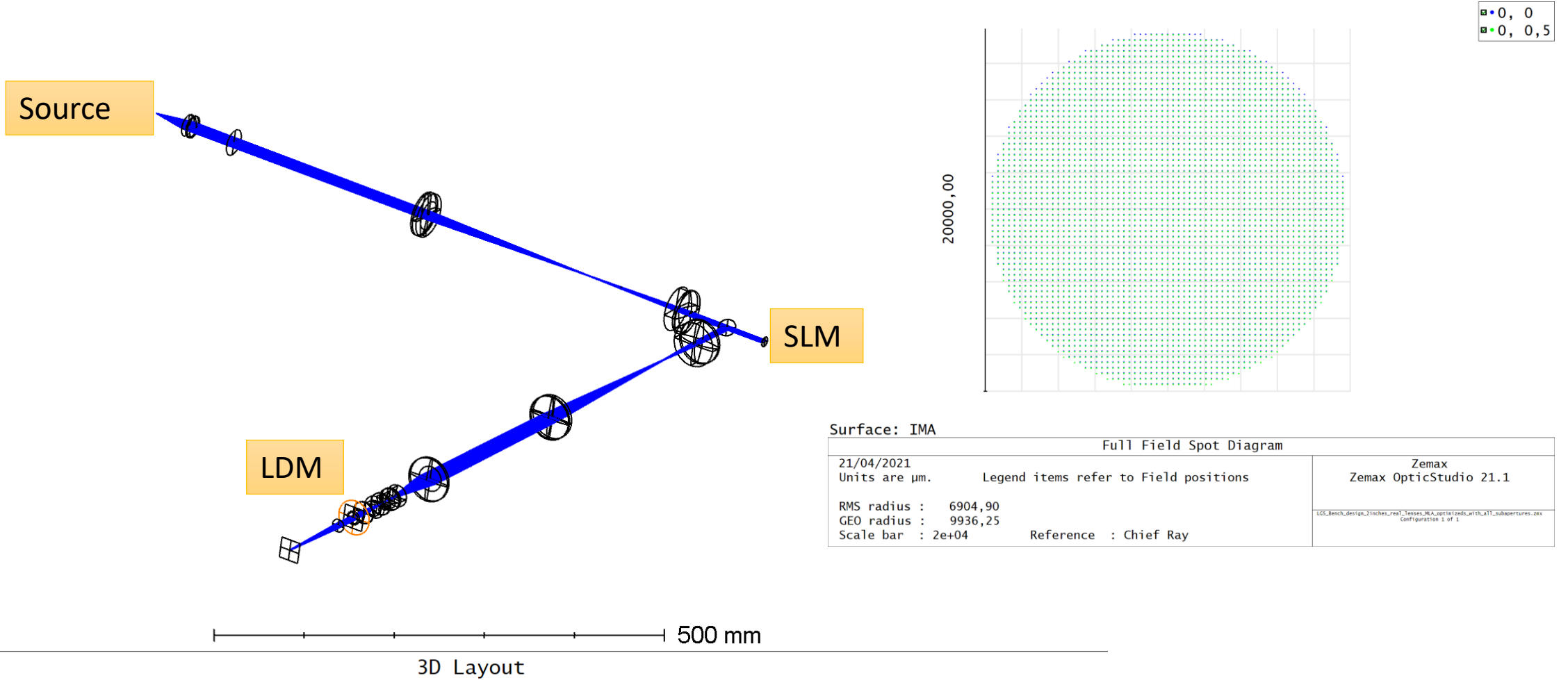


M4 deformable mirror (2,4m)



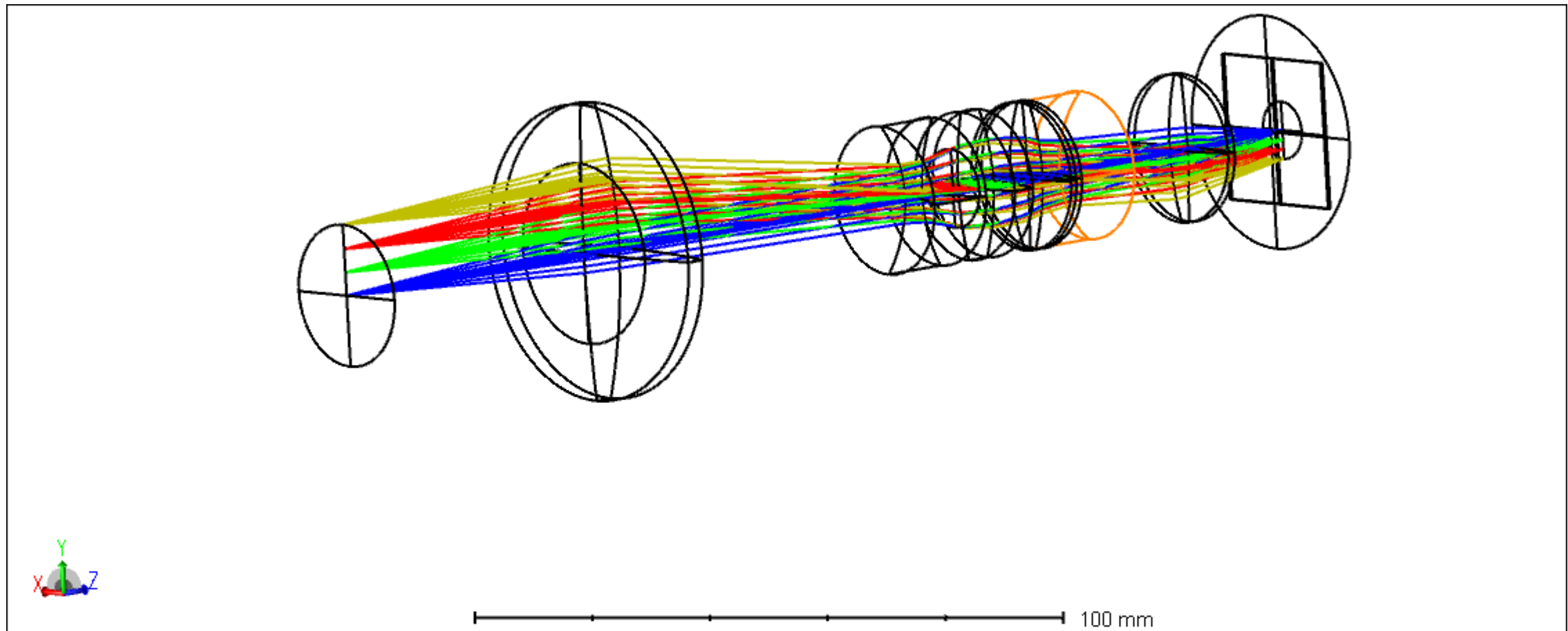
The sodium layer is around 90km high and its thickness is about 20km. The lasers are situated on the side of the telescope, therefore the wavefront sensor will see an elongated spot.

Optical Bench



Surface: IMA		Full Field Spot Diagram	
21/04/2021	Units are μm .	Legend items refer to Field positions	Zemax Zemax OpticStudio 21.1
RMS radius : 6904,90	GEO radius : 9936,25	Scale bar : 2e+04	Reference : Chief Ray
<small>LGS_Bench_design_2inches_real_lenses_and_relay.zmx Configuration 1 of 1</small>			

The LDM (Laser guide star detector module)



3D Layout

07/06/2021

Zemax
Zemax OpticStudio 21.1

HARMONI LDM RELAY PROTO SONY 2,48+6T.zmx
Configuration 1 of 1