

Provence Adaptive-optics PYramid RUn System

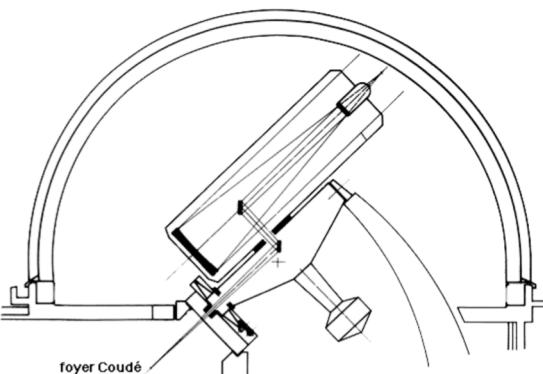
First on-sky results



E. Muslimov, N. Levraud, V. Chambouleyron, R. Félick, I. Boudjema, F. Pedreros, E. Soria-Hernandez, A. Lau, R. Lhoussaine, S. Karkar, A. Caillat, F. Leroux, M. Massa, M. Lopez, M. El Morsy, M. Cissé, F. Madec, K. El-Hadi, J-F Sauvage, B. Neichel, T. Fusco, J. Schmitt



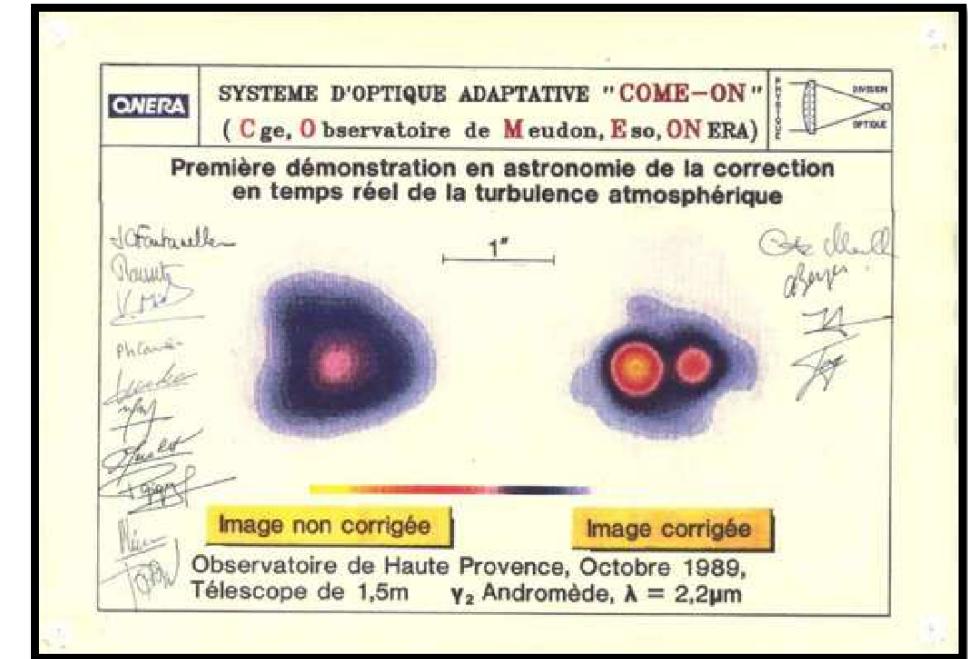
Adaptive optics system on the T-152 at OHP



T-152

- Coudé Focus
- f/D = 28
- Availability = 200 night / years
- Median seeing: 2"

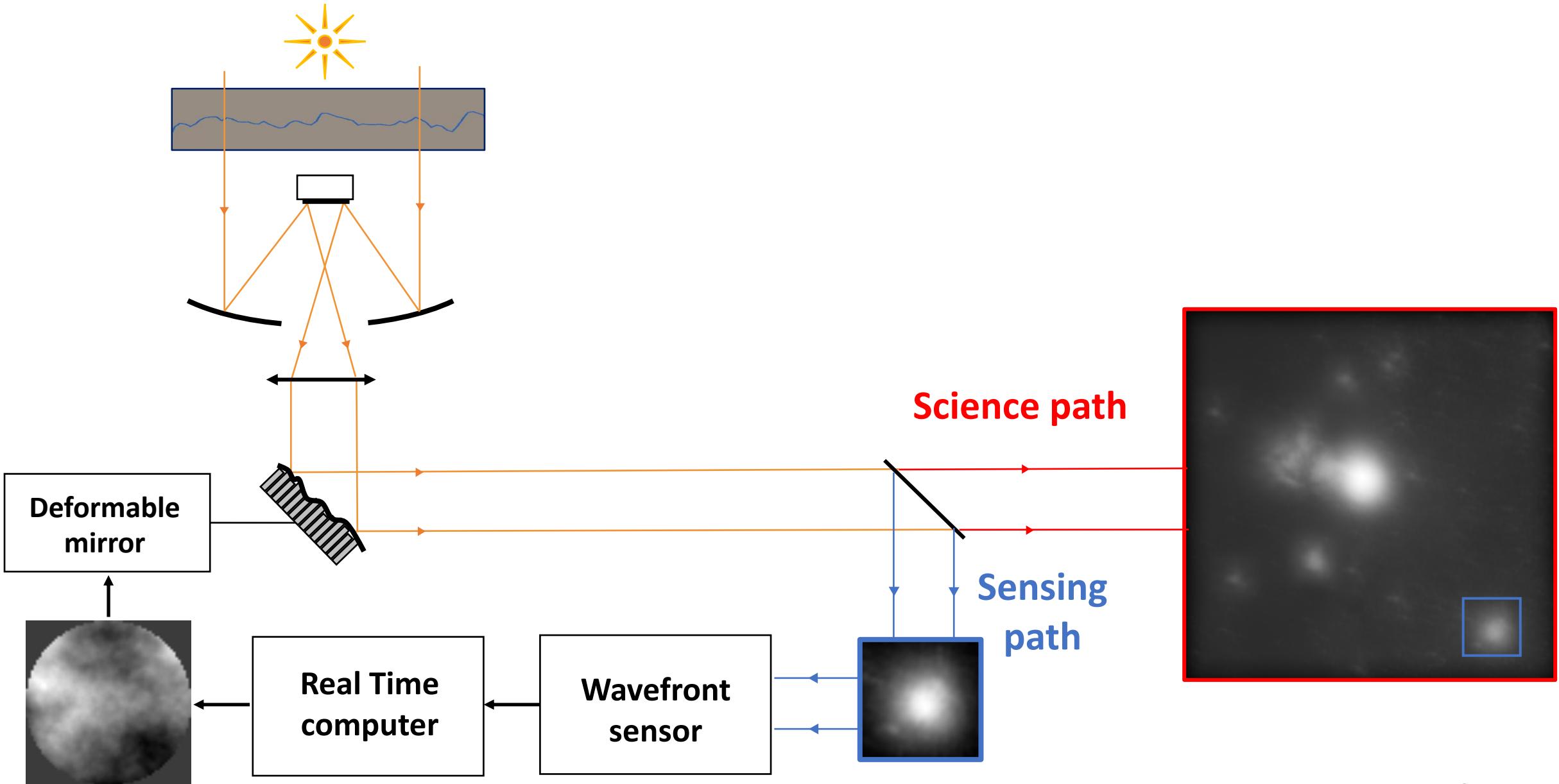
1989 : COME-ON



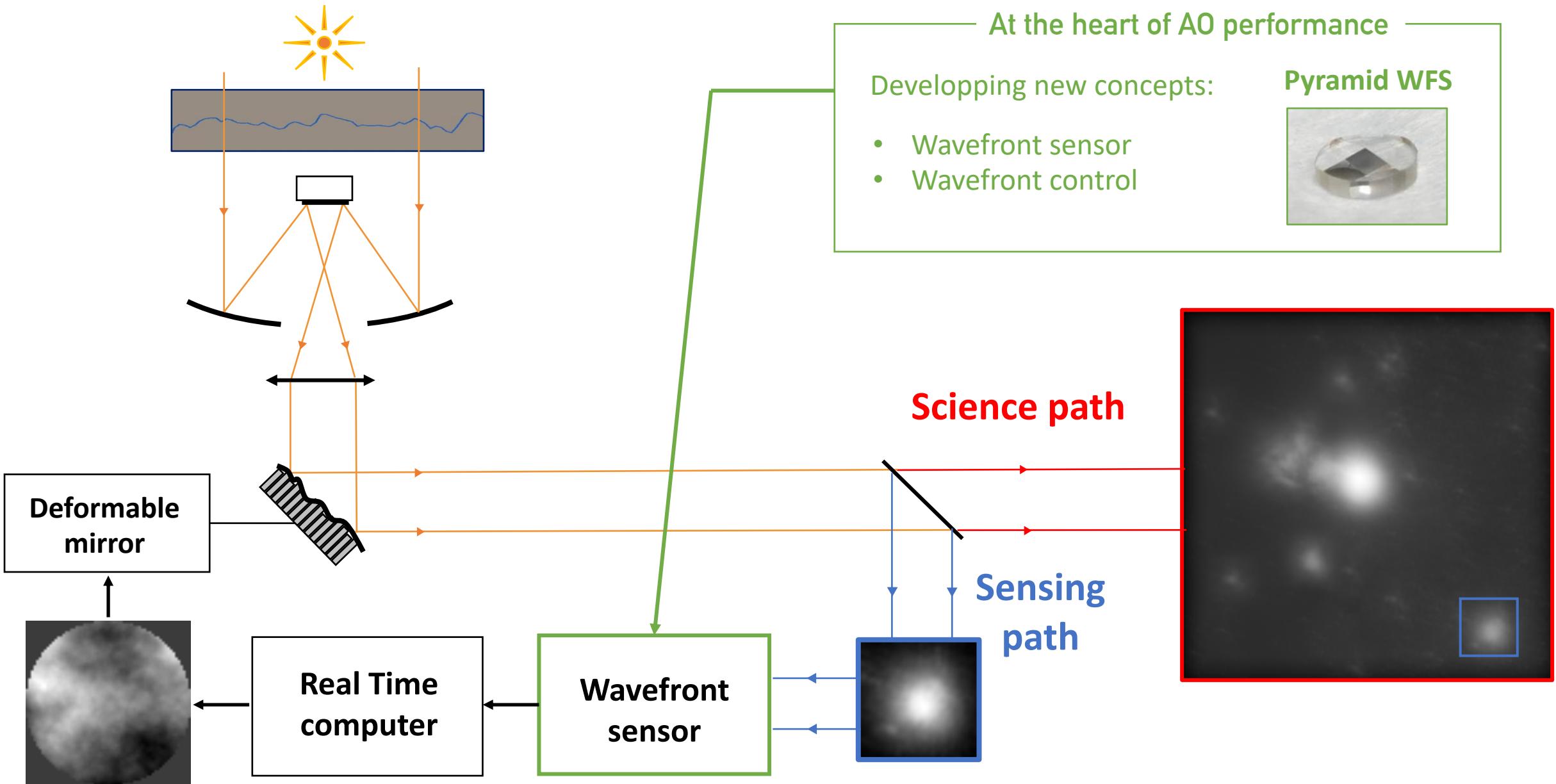
First astronomical AO corrected images !

19 actuators
5x5 Shack-Hartmann

A pyramid wavefront sensor on sky

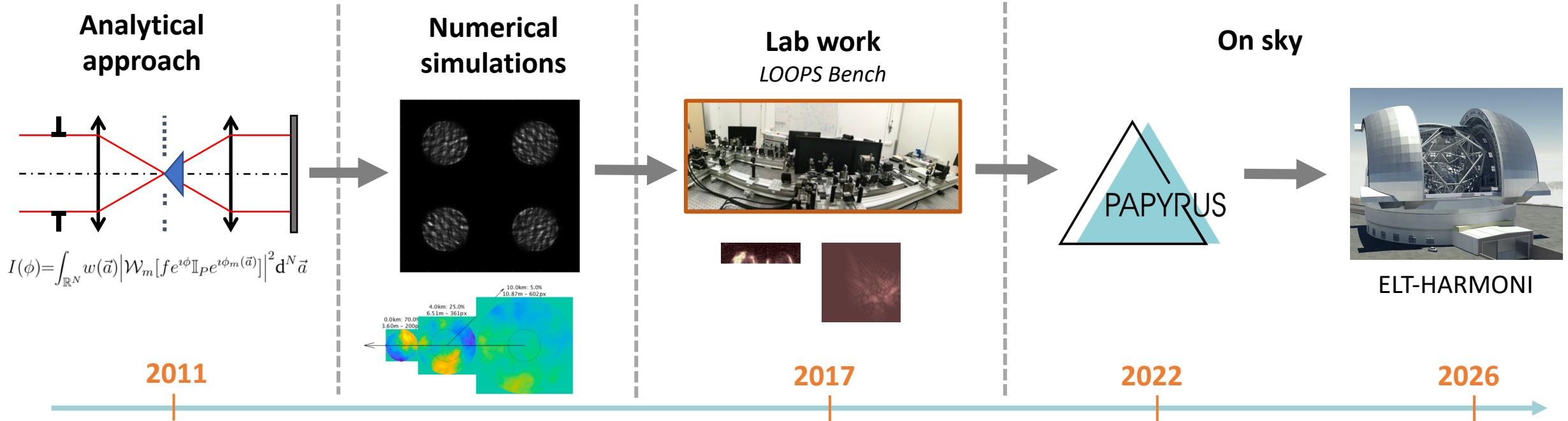


A pyramid wavefront sensor on sky



Instrumental R&D

- Test and demonstrate AO procedure with on-sky pyramid wavefront sensor
- Test new concepts of wavefront sensing and control
- Test new hardware : EMCCD camera, Deformable mirror, RTC, etc...



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Educational

- Project developped and led by students: Ph.D., interns and Postdocs (kick-off: January 2020)
- Pedagogic purpose: to be used for summer school at OHP

1st pedagogical Pyramid bench in the world

Astronomical

- High angular resolution back at OHP (80 mas in the visible)
- Possibility to add new instruments: coronograph, spectrometer, etc...
- A lot of available nights: **PAPYRUS will be soon open to community !**

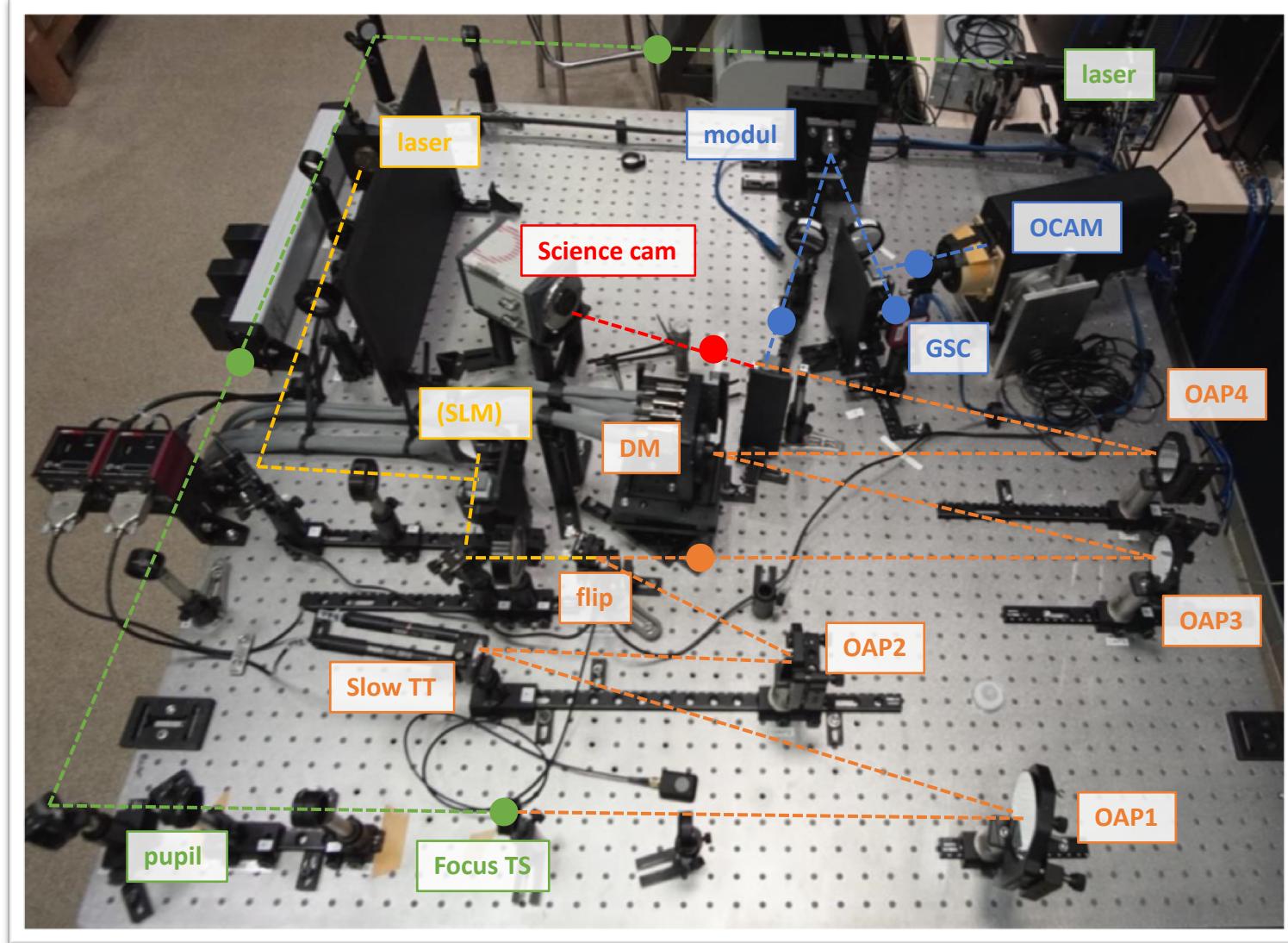
PAPYRUS optical layout



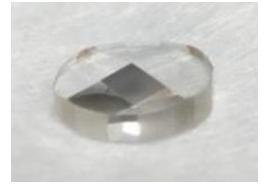
ALPAO - 17x17

Median $r_0(550 \text{ nm}) = 5 \text{ cm}$

1 actuator for $2 \times r_0$



120 cm



FLI - EMCCD

LEGEND

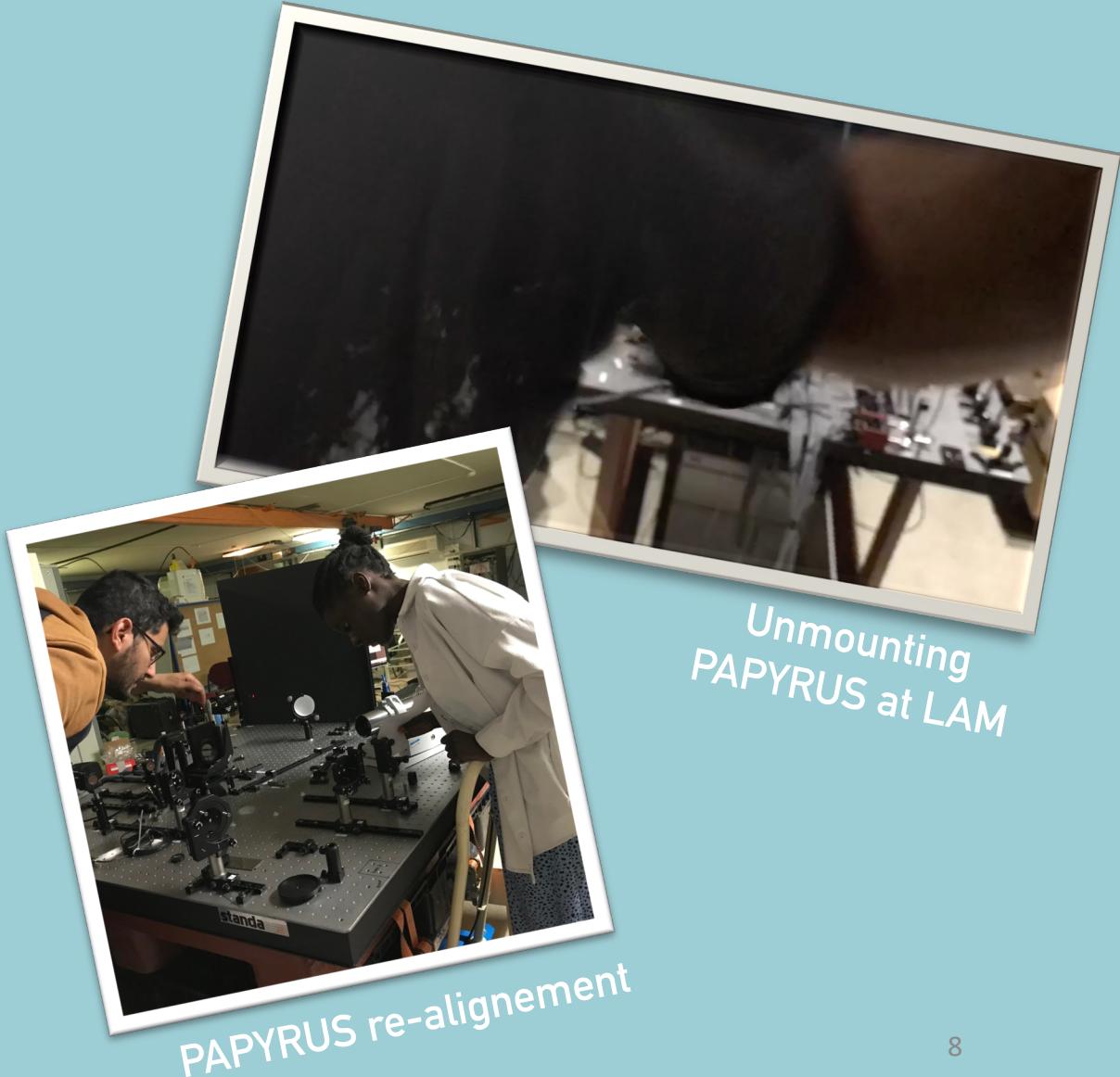
- Telescope simulator
- Common path
- Sensing path
- Science path
- Calibration unit



First light at OHP: June, 07th – June, 17th



PapYTEAM at the end of the mission



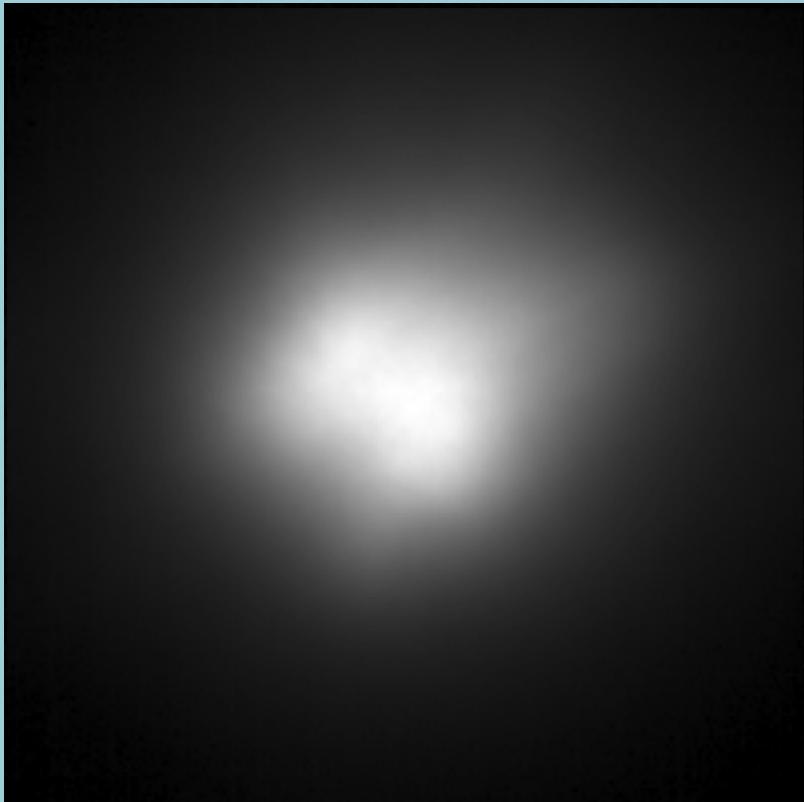
PAPYRUS re-alignement

Unmounting
PAPYRUS at LAM



First light at OHP: June, 07th – June, 17th

Strehl Ratio = 0.3%



Strehl Ratio = 15%



Véga

mag = 0

Loop frequency = 500 Hz

$\lambda = 635 \text{ nm} - \Delta\lambda = 40 \text{ nm}$

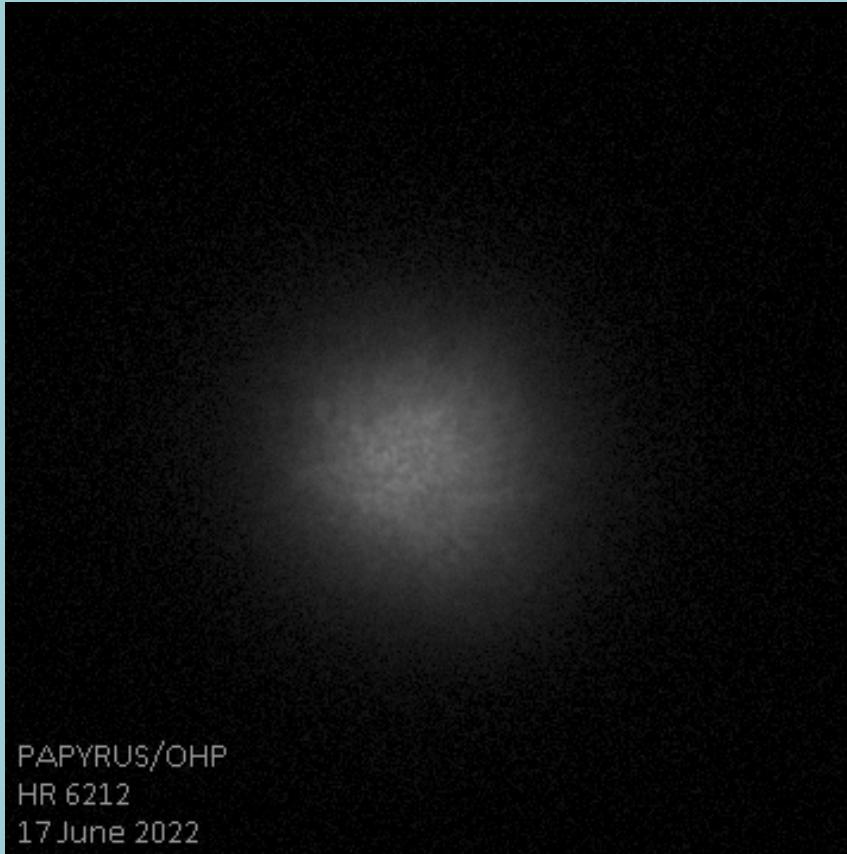


First light at OHP: June, 07th – June, 17th

Close loop on Binary star

HR 6212

- **mag1 = 2.95**
- **mag2 = 5.4**
- **separation = 1.6"**



Loop frequency = 500 Hz

$$\lambda = 635 \text{ nm} - \Delta\lambda = 40 \text{ nm}$$

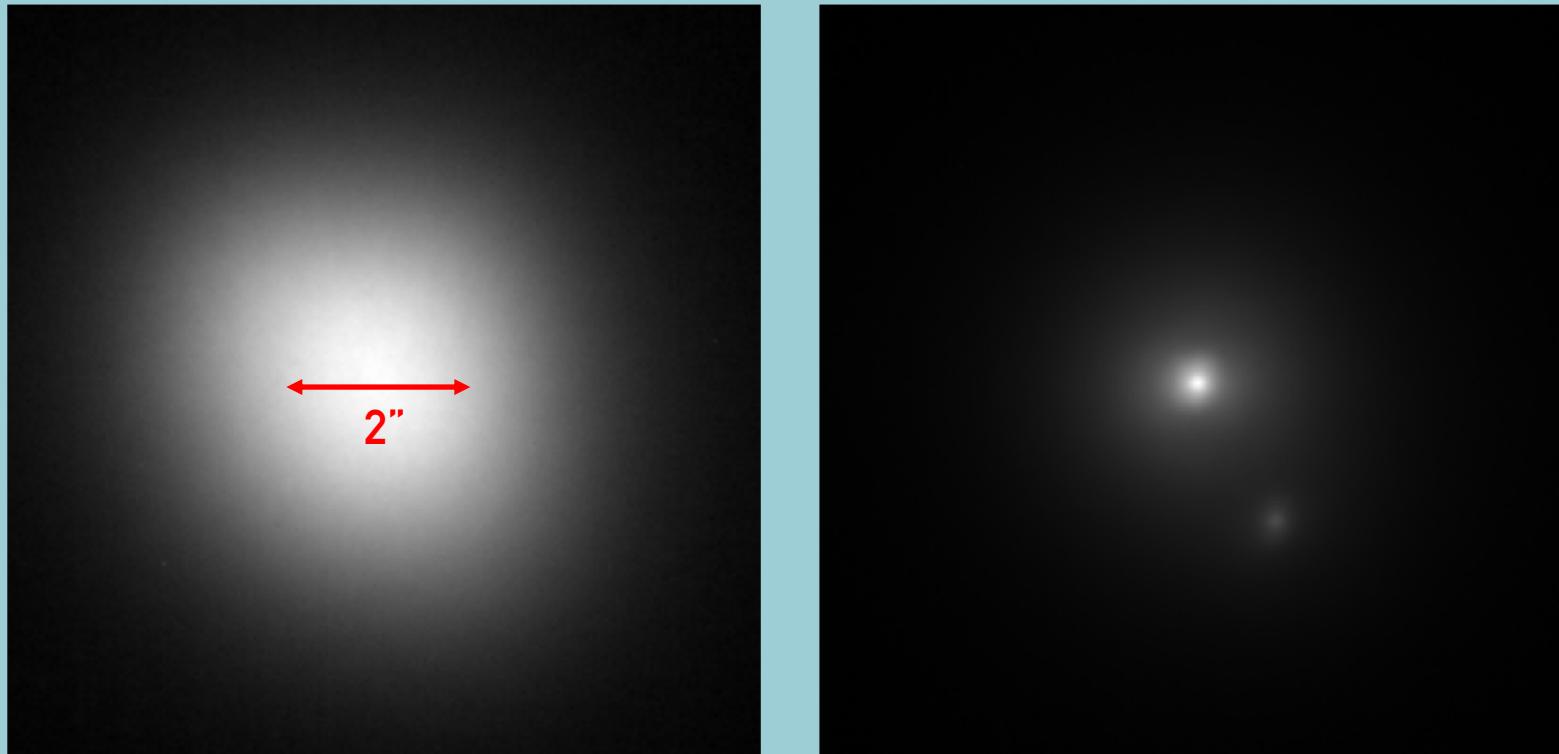


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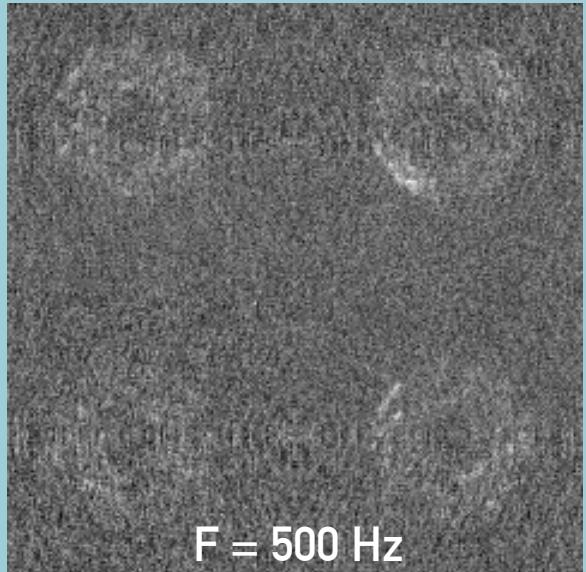


First light at OHP: June, 07th – June, 17th



Arcturus

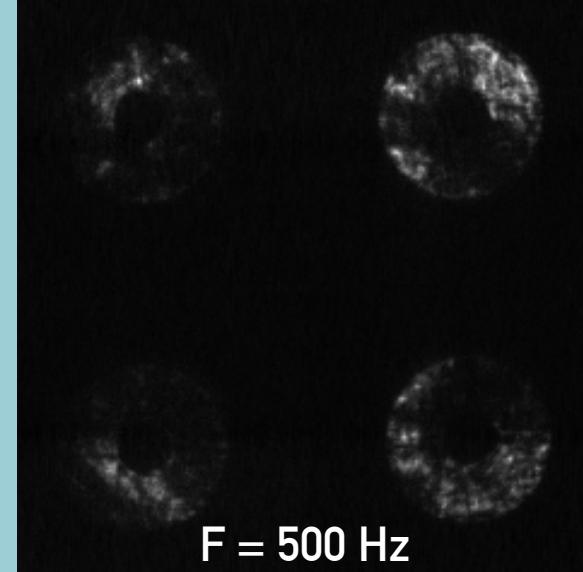
mag = 0.86



$F = 500 \text{ Hz}$



Electron
multiplying ON



$F = 500 \text{ Hz}$



Long exposure PSF

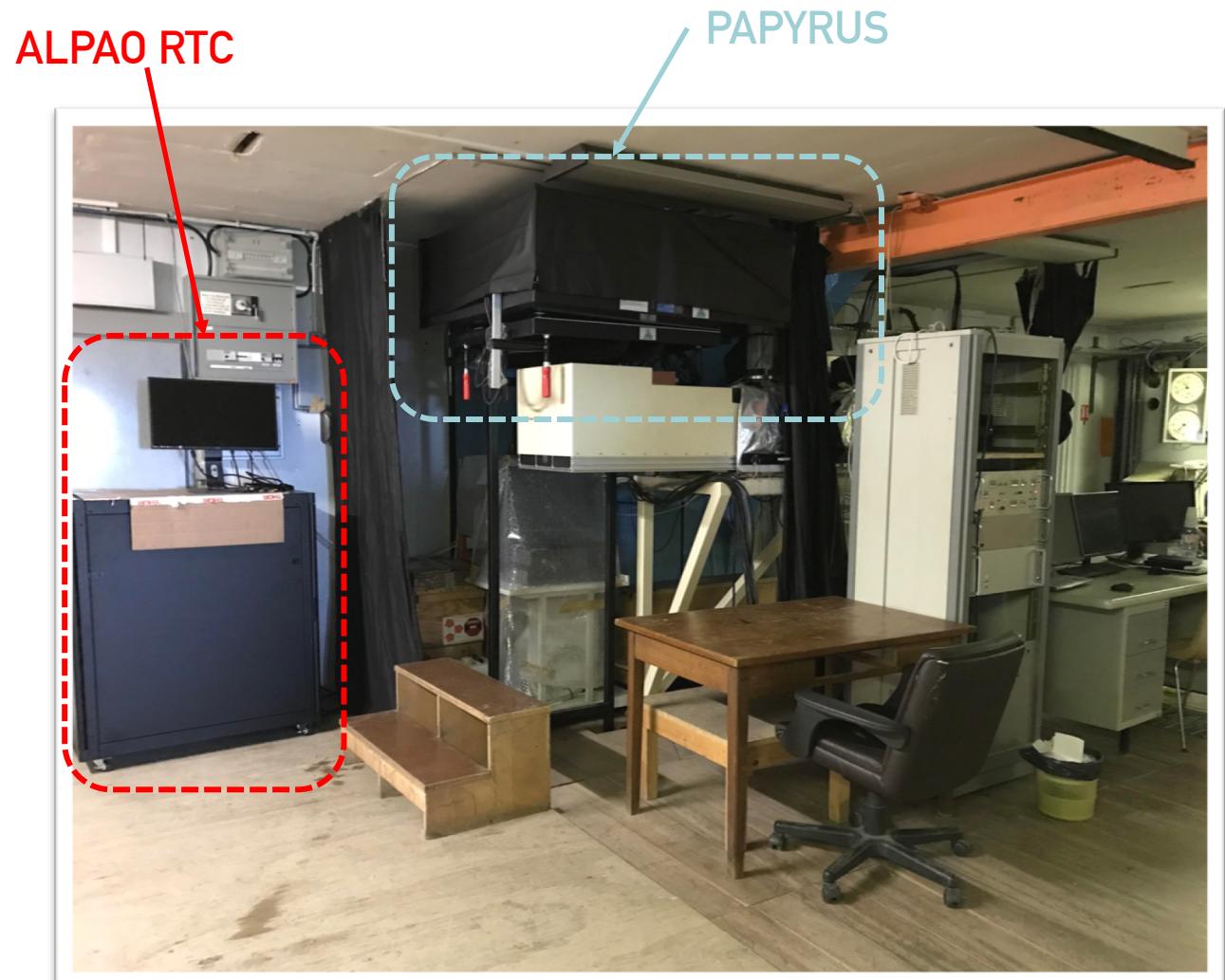


Short-term

- Increase temporal performance:
ALPAO RTC
- Turbulence generator (SLM) in
Calibration Unit
- Improve instrument throughput

Mid/long-term

- Test WFS control algorithms
developed at LAM
- Add WFS branch: Shack-Hartmann
- Welcome new science instruments





PAPYRUS

Thanks for listening